Introduction to black holes in two dimensions

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Outline



Gravity in 2D

- Models in 2D
- Generic dilaton gravity action

2 Classical and Semi-Classical Black Holes

- Classical black holes
- Semi-Classical black holes and thermodynamics
- 3 Quantum and Virtual Black Holes
 - Path integral quantization
 - S-matrix





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What is a black hole? Fishy analogy (Bill Unruh)





Above: black hole (NASA picture) Left: Waterfall

Analogy: Infinity +> Lake

Horizon ↔ Point of no return

Singularity ↔ Waterfall



2D or not 2D: that is the question

Riemann (Weyl+Ricci): $\frac{N^2(N^2-1)}{12}$ components in *N* dimensions

- 4D: 20 (10 Weyl and 10 Ricci)
- 5D: 50 (35 Weyl and 15 Ricci)
- 10D: 825 (770 Weyl and 55 Ricci)
- 11D: 1210 (1144 Weyl and 66 Ricci)
- 3D: 6 (Ricci)
- 2D: 1 (Ricci scalar) \rightarrow Lowest dimension with curvature
- 1D: 0

But: 2D EH $I_{EH} = \kappa \int d^2x \sqrt{gR}$: no equations of motion! Number of graviton modes: $\frac{N(N-3)}{2}$

Stuck already in the formulation of the model?

Have to go beyond Einstein-Hilbert in 2D!



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Example of 2D black hole

Schwarzschild black hole







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Artistic impression

Spherical symmetry

Carter-Penrose diagram of Schwarzschild

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- Spherical symmetry reduces 4D to 2D
- 2D: Time and surface radius
- Exact solution of Einstein equations: Schwarzschild
- Schwarzschild: "Hydrogen atom of General Relativity"
- Quantize in 2D!

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Spherical reduction

Line element adapted to spherical symmetry:

$$\mathrm{d} s^2 = \underbrace{g^{(N)}_{\mu\nu}}_{\mathrm{full \ metric}} \mathrm{d} x^{\mu} \, \mathrm{d} x^{\nu} = \underbrace{g_{\alpha\beta}(x^{\gamma})}_{2D \ \mathrm{metric}} \mathrm{d} x^{\alpha} \, \mathrm{d} x^{\beta} - \underbrace{\phi^2(x^{\alpha})}_{\mathrm{surface \ area}} \mathrm{d} \Omega^2_{\mathrm{S}_{N-2}} \,,$$

Insert into N-dimensional EH action $I_{EH} = \kappa \int d^N x \sqrt{-g^{(N)}} R^{(N)}$:



Cosmetic redefinition $X \propto (\lambda \phi)^{N-2}$:

$$I_{EH} \propto \int d^2 x \sqrt{-g} \left[\underbrace{XR + \frac{N-3}{(N-2)X} (\nabla X)^2 - \lambda^2 X^{(N-4)/(N-2)}}_{\text{Scalar-tensor theory a.k.a. dilaton gravity}} \right]_{\text{INTECORE ACTIONS}}$$

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Second order formulation

Similar action arises from string theory, from other kinds of dimensional reduction, from intrinsically 2D considerations, ... Generic action:

$$I_{2DG} = \kappa \int d^2 x \sqrt{-g} \Big[XR + U(X)(\nabla X)^2 - V(X) \Big]$$
(1)

Special case U = 0, $V = X^2$: EOM R = 2X

$$I\propto\int\mathrm{d}^2x\sqrt{-g}R^2$$

Similarly f(R) Lagrangians related to (1) with U = 0String context: $X = e^{-2\phi}$, with ϕ as string dilaton Conformal trafo to different model with $\tilde{U}(X) = 0$: $\tilde{V}(X) = \frac{d}{dX} \underbrace{w(X) := V(X)e^{Q(X)}}_{W(X)}$, with $Q(X) := \int^X dy U(y)$

conformally invariant



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Selected list of models

Model	U(X)	V(X)	w(X)
1. Schwarzschild (1916)	$-\frac{1}{2X}$	$-\lambda^2$	$-2\lambda^2\sqrt{X}$
2. Jackiw-Teitelboim (1984)	0	٨X	$\frac{1}{2}\Lambda X^2$
3. Witten BH (1991)	$-\frac{1}{2}$	$-2b^2X$	$-2b^2X$
4. CGHS (1992)	o î	$-2b^{2}$	$-2b^2X$
5. (A)dS ₂ ground state (1994)	$-\frac{a}{v}$	BX	$a \neq 2 : \frac{B}{2a} X^{2-a}$
6. Rindler ground state (1996)	$-\frac{\dot{a}}{\nabla}$	BX ^a	BX
7. BH attractor (2003)	o ô	BX ⁻¹	B In X
8. SRG (N > 3)	$-\frac{N-3}{(N-2)X}$	$-\lambda^2 X^{(N-4)/(N-2)}$	$-\lambda^2 \frac{N-2}{N-3} X^{(N-3)/(N-2)}$
9. All above: ab-family (1997)	$-\frac{a}{X}$	BX ^{a+b}	$b \neq -1$: $\frac{B}{b+1}X^{b+1}$
10. Liouville gravity	а	be $^{\alpha X}$	$a \neq -\alpha$: $\frac{b}{a+\alpha} e^{(a+\alpha)X}$
11. Reissner-Nordström (1916)	$-\frac{1}{2X}$	$-\lambda^2 + \frac{Q^2}{X}$	$-2\lambda^2\sqrt{X}-2Q^2/\sqrt{X}$
12. Schwarzschild-(A)dS	$-\frac{1}{2X}$	$-\lambda^2 - \hat{\ell X}$	$-2\lambda^2\sqrt{X}-\tfrac{2}{3}\ell X^{3/2}$
13. Katanaev-Volovich (1986)	α	$\beta X^2 - \Lambda$	$\int^X e^{\alpha y} (\beta y^2 - \Lambda) dy$
14. Achucarro-Ortiz (1993)	0	$\frac{Q^2}{X} - \frac{J}{4X^3} - \Lambda X$	$Q^2 \ln X + \frac{J}{R \chi^2} - \frac{1}{2} \Lambda X^2$
15. Scattering trivial (2001)	generic	Ô	const.
16. KK reduced CS (2003)	0	$\frac{1}{2}X(c-X^2)$	$-\frac{1}{8}(c-X^2)^2$
17. exact string BH (2005)	lengthy	$-\gamma$	$-(1 + \sqrt{1 + \gamma^2})$
18. Symmetric kink (2005)	generic	$-X\Pi_{i=1}^{n}(X^{2}-X_{i}^{2})$	lengthy
19. KK red. conf. flat (2006)	$-\frac{1}{2} \tanh{(X/2)}$	A sinh X	4A cosh (X/2)
20. 2D type 0A	$-\frac{1}{X}$	$-2b^2X+\frac{b^2q^2}{8\pi}$	$-2b^2X+rac{b^2q^2}{8\pi}\ln X$

Red: mentioned in abstract

Blue: pioneer models

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Daniel Grumiller Introduction to black holes in two dimensions

First order formulation

Gravity as gauge theory (89: Isler, Trugenberger, Chamseddine, Wyler, 91: Verlinde, 92: Cangemi, Jackiw, Achucarro, 93: Ikeda, Izawa, 94: Schaller, Strobl)

Example: Jackiw-Teitelboim model ($U = 0, V = \Lambda X$)

$$[P_a, P_b] = \Lambda \varepsilon_{ab} J, \qquad [P_a, J] = \varepsilon_a{}^b P_b,$$

Non-abelian BF theory:

$$I_{BF} = \int X_{A}F^{A} = \int \left[X_{a} de^{a} + X_{a} \varepsilon^{a}{}_{b} \omega \wedge e^{b} + X d\omega + \varepsilon_{ab} e^{a} \wedge e^{b} \Lambda X \right]$$

field strength F = dA + [A, A]/2 contains SO(1, 2) connection $A = e^a P_a + \omega J$, coadjoint Lagrange multipliers X_A Generic first order action:



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$$I_{2DG} \propto \int \left[X_a \underbrace{T^a}_{\text{torsion}} + X \underbrace{R}_{\text{curvature}} + \underbrace{\epsilon}_{\text{volume}} (X^a X_a U(X) + V(X)) \right]$$
(2)
$$T^a = de^a + \varepsilon^a{}_b \omega \wedge e^b, R = d\omega, \epsilon = \varepsilon_{ab} e^a \wedge e^b_{\text{curvature}}$$

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Classical solutions

Light-cone components and Eddington-Finkelstein gauge (87: Polyakov, 92: Kummer, Schwarz, 96: Klösch, Strobl)

Constant dilaton vacua:

$$X = \text{const.}, \quad V(X) = 0, \quad R = V'(X)$$

- Minkowski, Rindler or (A)dS only
- isolated solutions (no constant of motion)

Generic solutions in EF gauge $\omega_0 = e_0^+ = 0$, $e_0^- = 1$:

$$ds^{2} = 2e^{Q(X)} du dX + \underbrace{e^{Q(X)}(w(X) + M)}_{Killing \text{ scent}} du^{2}$$

• Birkhoff theorem: at least one Killing vector ∂_u

- one constant of motion: mass M
- dilaton is coordinate x⁰ (residual gauge trafos!)

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(3)

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Global structure

Simple algorithm exists to construct all possible global structures (Israel, Walker)



Key ingredient: Killing norm!

- for each zero w(X) + M = 0: Killing horizon
- multiple zeros: extremal horizons (BPS)
- glue together basic EF-patches
- caveat: bifurcation points
- check geodesics for (in)completeness
- Simple example: Carter-Penrose diagram on the left: Killing norm $1 2M/r + Q^2/r^2$ (RN)



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Hawking radiation Quantization on fixed background; method by Christensen and Fulling



• conformal anomaly $< T^{\mu}{}_{\mu} > \propto R$

- conservation equation $abla_{\mu} < T^{\mu
 u} >= 0$
- boundary conditions (Unruh, Hartle-Hawking, Boulware)

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• get flux from trace

2D Stefan-Boltzmann (flux $\propto T_H^2$):

$$T_H = \frac{1}{2\pi} |w'(X)|_{X=X_h} = \text{surface gravity}$$

other thermodynamical quantities of interest: entropy: X on horizon, specific heat: w'/w'' on horizon



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Partition function:

$$Z = \int \mathcal{D}g\mathcal{D}X e^{-rac{1}{\hbar} \Gamma[g_{\mu
u},X]}$$

Euclidean action:

 $\mathsf{F}[g_{\mu\nu}, X] = I_{\text{bulk}}[g_{\mu\nu}, X] + I_{\text{GHY}}[\gamma_{\mu\nu}, X] + I_{\text{counter}}[\det(\gamma_{\mu\nu}), X]$

$$\Gamma[p,q] = \underbrace{\int dt[-q\dot{p} - H(p,q)]}_{\text{bulk term}} + \underbrace{qp|_{t_i}^{t_f}}_{\text{Gibbons-Hawking-York}} + \underbrace{C(q)|_{t_i}^{t_f}}_{\text{counter term}}$$

- Bulk term: "usual" action
- GHY: boundary conditions
- Counter term: consistency of path integral

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- toy models (2D strings, black hole evaporation, ...)
- classically: spherically symmetric sector of general relativity (critical collapse)
- semi-classically: near horizon geometry effectively 2D (Carlip, Wilczek, ...)
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- more speculative ideas: at high energies gravity effectively 2D (Reuter, Ambjorn, Loll)? gravity near the Earth: linear potential, i.e., effectively 2D (Mann, Young)?



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Daniel Grumiller Introduction to black holes in two dimensions

Outline

Gravity in 2D

- Models in 2D
- Generic dilaton gravity action
- 2 Classical and Semi-Classical Black Holes
 - Classical black holes
 - Semi-Classical black holes and thermodynamics
- Quantum and Virtual Black HolesPath integral quantization
 - S-matrix



Quantization of specific models

Non-comprehensive history

- 1992: Cangemi, Jackiw (CGHS)
- 1994: Louis-Martinez, Gegenberg, Kunstatter (U = 0)
- 1994: Kuchař (Schwarzschild)
- 1995: Cangemi, Jackiw, Zwiebach (CGHS)
- 1997: Kummer, Liebl, Vassilevich (generic geometry)
- 1999: Kummer, Liebl, Vassilevich (minimally coupled scalar, generic geometry)
- 2000-2002: DG, Kummer, Vassilevich (non-minimally coupled scalar, generic geometry)
- 2004: Bergamin, DG, Kummer (minimally coupled matter, generic SUGRA)
- 2006: DG, Meyer (non-minimally coupled fermions, generic geometry)



- no matter: integrability, no scattering, no propagating physical modes
- with matter: no integrability in general, scattering, critical collapse

Massless scalar field S:

$$I_m = \int \mathrm{d}^2 x \sqrt{-g} F(X) (\nabla S)^2$$

- minimal coupling: F = const.
- non-minimal coupling otherwise
- spherical reduction: $F \propto X$

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Non-perturbative path integral quantization

- constraint analysis $\{G^{i}(x), G^{j}(x')\} = G^{k}C_{k}{}^{ij}\delta(x-x')$
- BRST charge $\Omega = c^i G_i + c^i c^j C_{ij}^k p_k$ (ghosts c^i, p_k)
- gauge fixing fermion to achieve EF gauge

integrating ghost sector yields

$$Z[\text{sources}] = \int \mathcal{D}f\delta\left(f + i\delta/\delta j_{\theta_1^+}\right) \tilde{Z}[f, \text{sources}]$$

with $(\tilde{S} = S\sqrt{f})$

 $\tilde{Z}[f, \text{sources}] = \int \mathcal{D}\tilde{S}\mathcal{D}(\omega, e^a, X, X^a) \det \Delta_{F.P.} \exp i(I_{g.f.} + \text{sources})$

Can integrate over all fields except matter non-perturbatively!



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Non-local effective theory

Convert local gravity theory with matter into non-local matter theory without gravity

Generating functional for Green functions (F = 1):

$$\tilde{Z}[f, \text{sources}] = \int \mathcal{D}\tilde{S} \exp i \int (\mathcal{L}^k + \mathcal{L}^v + \mathcal{L}^s) d^2x$$

 $\mathcal{L}^{k} = \partial_{0} S \partial_{1} S - E_{1}^{-} (\partial_{0} S)^{2}, \ \mathcal{L}^{v} = -w'(\hat{X}), \ \mathcal{L}^{s} = \sigma S + j_{e^{+}} \hat{E}_{1}^{+} + \dots,$

$$\tilde{S} = Sf^{1/2}, \ \hat{E}_1^+ = e^{Q(\hat{X})}, \ \hat{X} = \underbrace{a+bx^0}_{X} + \underbrace{\partial_0^{-2}(\partial_0 S)^2}_{\text{non-local}} + \dots, \ a = 0, \ b = 1,$$

$$\int \mathcal{D}\tilde{S} \exp i \int \mathcal{L}^k = \exp\left(\frac{i}{96\pi} \int_X \int_Y fR_X \Box_{xy}^{-1} R_y\right)$$

Red: geometry, Magenta: matter, Blue: boundary, conditions, and an anter

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$$E_1^- = w(X) + M, \quad \hat{E}_1^+ = e^{Q(X)} + e^{Q(X)}U(X)\partial_0^{-2}(\partial_0 S)^2 + \dots$$

$$\int \mathcal{D}\tilde{\mathsf{S}} \exp i \int \mathcal{L}^{k} = \exp \underbrace{\left(i/96\pi \int_{x} \int_{y} f R_{x} \Box_{xy}^{-1} R_{y}\right)}_{\text{Polyakov}}$$

Red: geometry, Magenta: matter, Blue: boundary conditions

Some Feynman diagrams



- so far: calculated only lowest order vertices and propagator corrections
- partial resummations possible (similar to Bethe-Salpeter)?
- non-local loops vanish to this order

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S-matrix for s-wave gravitational scattering

Quantizing the Einstein-massless-Klein-Gordon model

ingoing s-waves $q = \alpha E$, $q' = (1 - \alpha)E$ interact and scatter into outgoing s-waves $k = \beta E, k' = (1 - \beta)E$

$$\tilde{T} = \Pi \ln \frac{\Pi^2}{E^6} + \frac{1}{\Pi} \sum_{p} p^2 \ln \frac{p^2}{E^2} \cdot \left(3kk' qq' - \frac{1}{2} \sum_{r \neq p} \sum_{s \neq r, p} r^2 s^2 \right)$$



- result finite and simple
- monomial scaling with E
- forward scattering poles $\Pi = 0$
- decay of s-waves possible



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> $T(q, q'; k, k') \propto \tilde{T} \delta(k + k' - q - q') / |kk'qq'|^{3/2}$ (4a)

with $\Pi = (k + k')(k - q)(k' - q)$ and

$$\tilde{T} = \Pi \ln \frac{\Pi^2}{E^6} + \frac{1}{\Pi} \sum_{p} p^2 \ln \frac{p^2}{E^2} \cdot \left(3kk' qq' - \frac{1}{2} \sum_{r \neq p} \sum_{s \neq r, p} r^2 s^2 \right)$$
(4b)



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Virtual black holes

Reconstruct geometry from matter



- Schwarzschild and Rindler terms
- nontrivial part localized
- geometry is non-local (depends on r, u, r_0, u_0)
- geometry asymptotically fixed (Minkowski)

- J. D. Brown, "LOWER DIMENSIONAL GRAVITY," World Scientific Singapore (1988).
- A. Strominger, "Les Houches lectures on black holes," hep-th/9501071.
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