

# Gravitational waves

A new tool for observing the Universe through ripples in spacetime

Daniel Grumiller

Institute for Theoretical Physics  
TU Wien

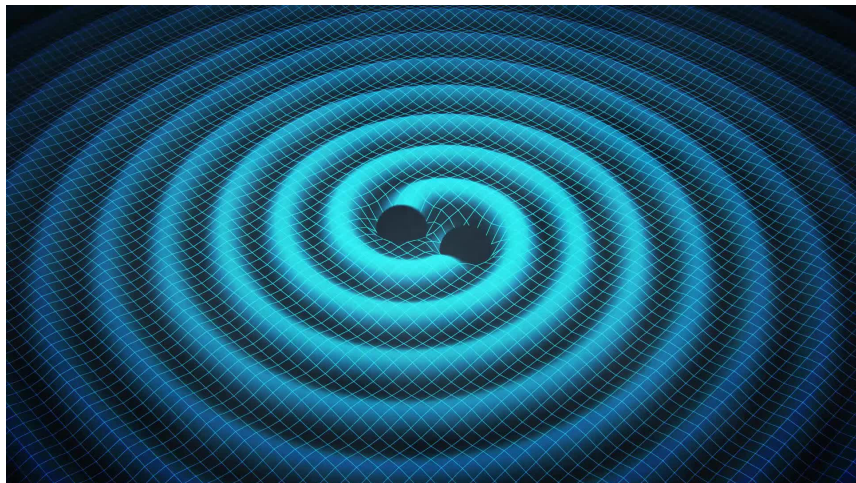
Based on public lecture for MIT Club Austria  
Black Holes I, January 2018



## Executive summary

- ▶ 1915: Einstein's General Relativity predicts gravitational waves

$$R_{\mu\nu} = 0$$



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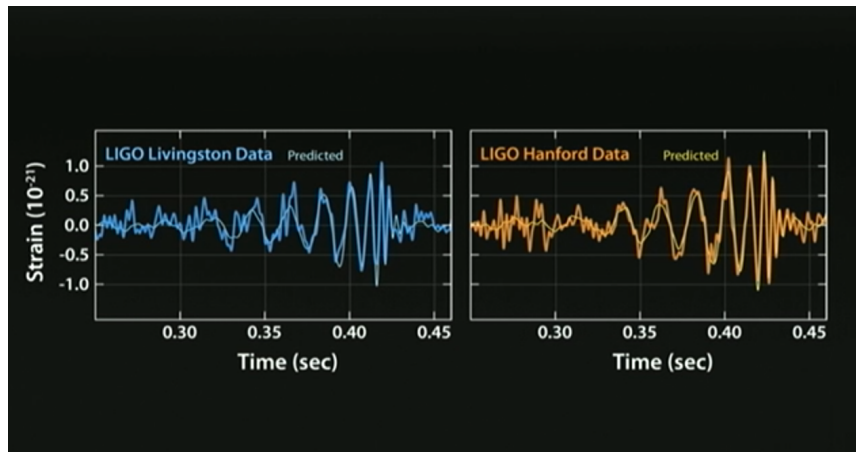
- ▶ 1915: Einstein's General Relativity predicts gravitational waves
- ▶ 1916: Schwarzschild constructs first black hole solution

$$ds^2 = -(1 - 2M/r) dt^2 + \frac{dr^2}{1 - 2M/r} + r^2 d^2\Omega_{S^2}$$



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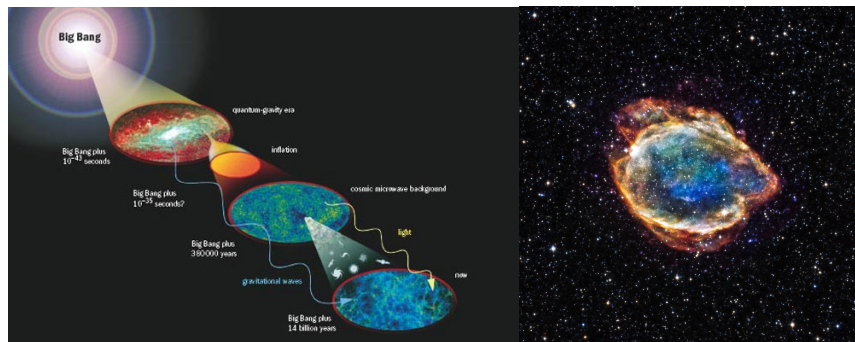
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- ▶ 2115: gravitational waves standard tool in astrophysics & cosmology



Left: gravitational waves sensitive to early Universe, Right: Supernova

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- ▶ 2116: someone pays taxes for gravitational waves



Disclaimer: quote above is commonly cited, but probably not authentic

# Outline

Theory

Technology

Observation

Applications



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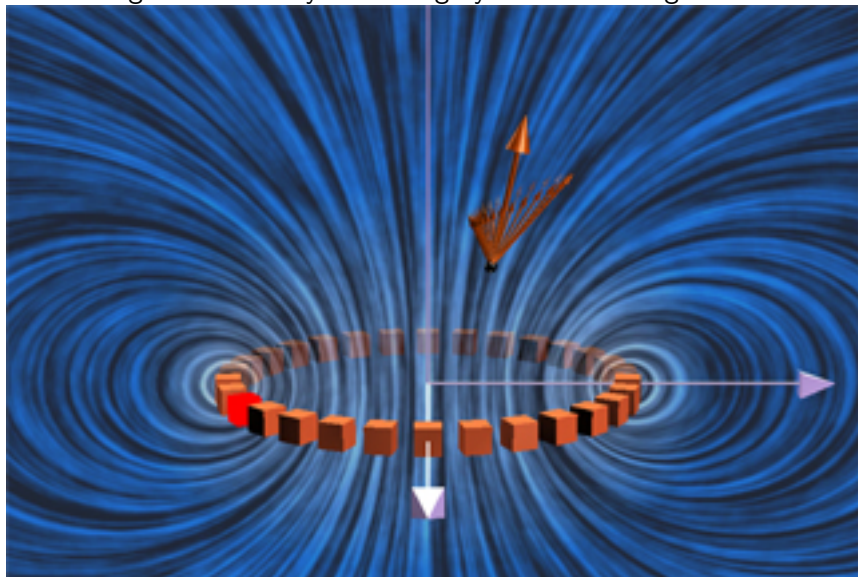
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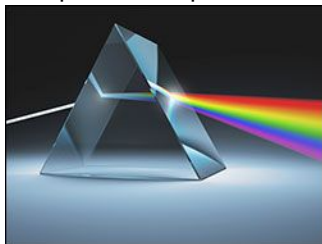
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- ▶ Electromagnetism: theory describing dynamics of charges

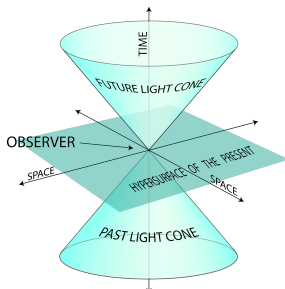


## Warm-up: electromagnetism

- ▶ Electromagnetism: theory describing dynamics of charges
- ▶ Unifies electricity, magnetism, optics and special relativity



$$E = mc^2$$



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- ▶ Electromagnetism: theory describing dynamics of charges
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- ▶ Quantum electrodynamics tested with amazing precision and accuracy

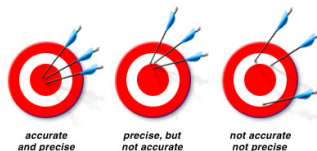
Example: gyromagnetic factor of electron

Experiment (2008):

$$\frac{g_e^{\text{exp}}}{2} = 1.00115965218073 \pm 0.000000000000028$$

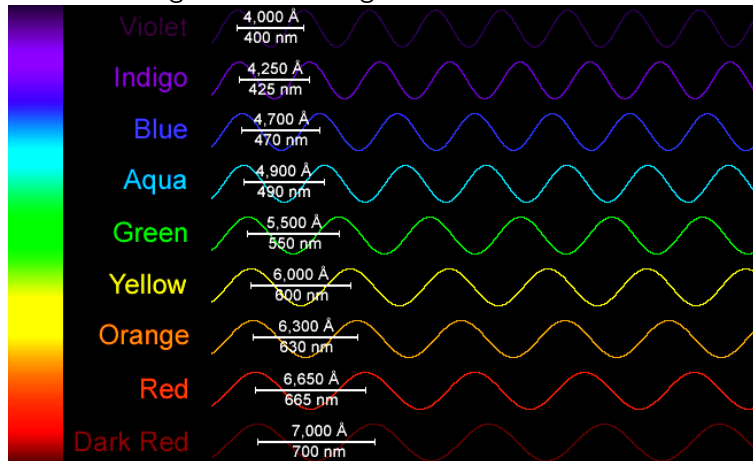
Theory (2012):

$$\frac{g_e^{\text{the}}}{2} = 1.00115965218178 \pm 0.000000000000077$$



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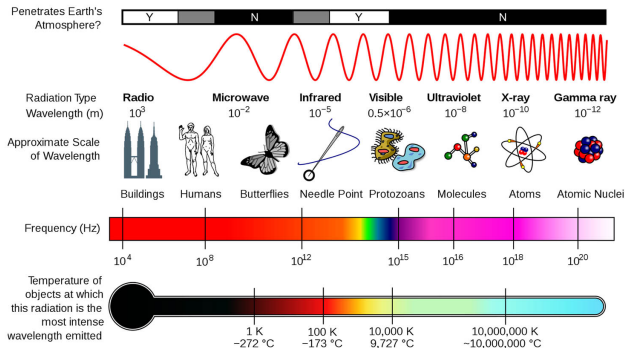
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$$dF = 0 = d * F \quad \Rightarrow \quad \square A_\mu = 0$$

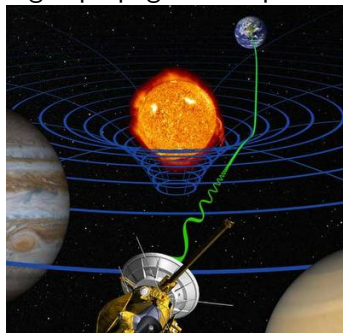


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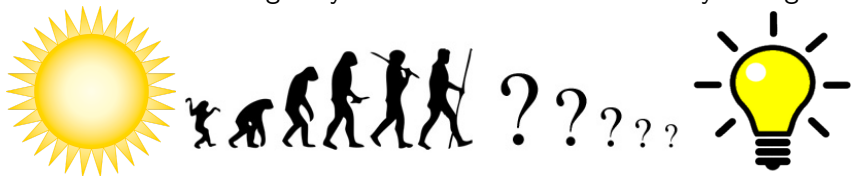


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- ▶ First observation of light by humankind: about 2 million years ago



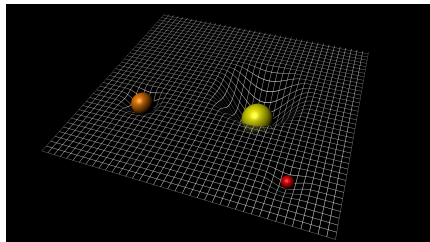
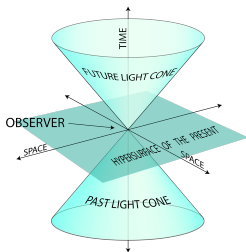
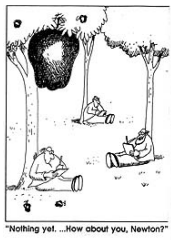
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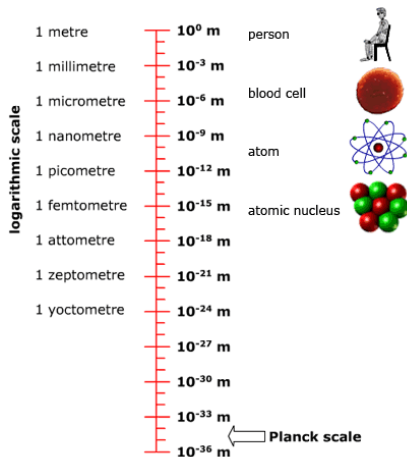
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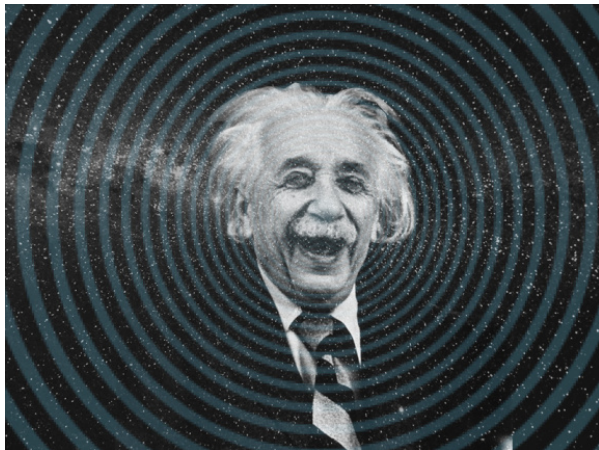
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[www.phys.unsw.edu.au/einsteinlight](http://www.phys.unsw.edu.au/einsteinlight)

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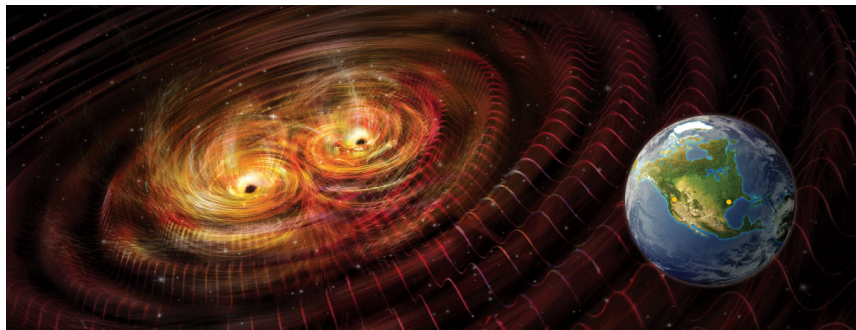
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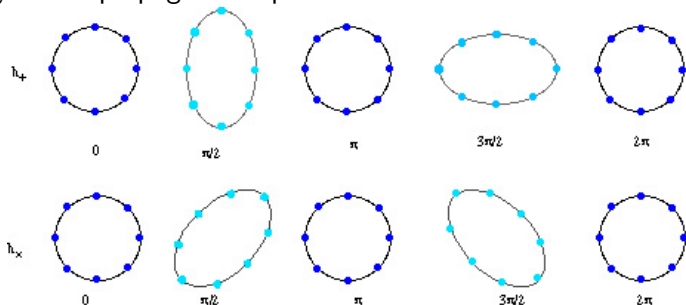


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## Gravitational wave sources

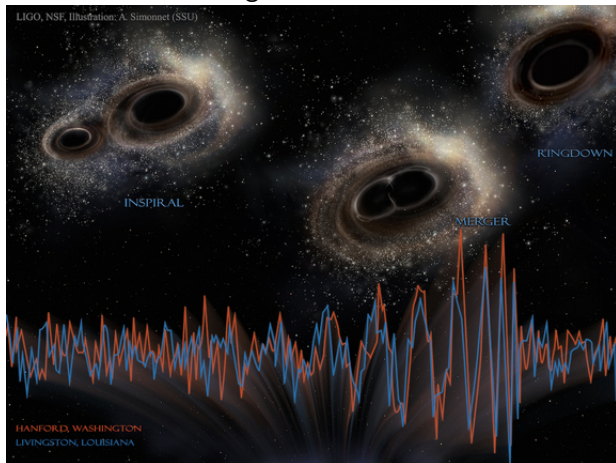
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- ▶ Simple calculation to estimate gravitational wave energy

Take two Schwarzschild black holes of equal mass  $m$

Call final black hole mass  $M$  and gravitational wave energy  $E$

Energy conservation:

$$M + E = m + m \Rightarrow E = 2m - M$$

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Get upper bound on gravitational wave energy

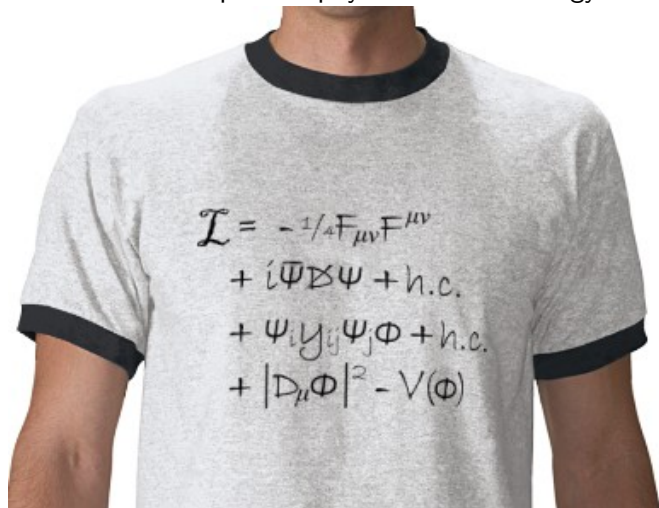
$$E \leq (2 - \sqrt{2})m \approx 29\% \text{ of initial energy}$$

Energy released by  $10^{34} - 10^{36}$  Nagasaki bombs!

## Gravitational waves as the last known puzzle piece

Big questions: what is the Universe made of? what are the fundamental forces?

- ▶ All known forces of Nature described with amazing precision through the Standard Models of particle physics and Cosmology

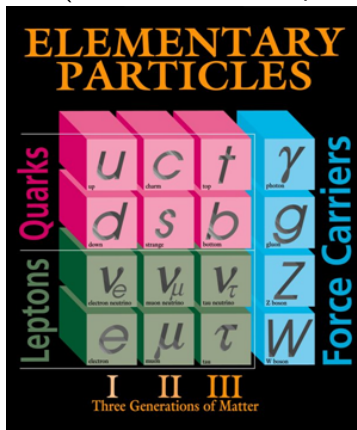




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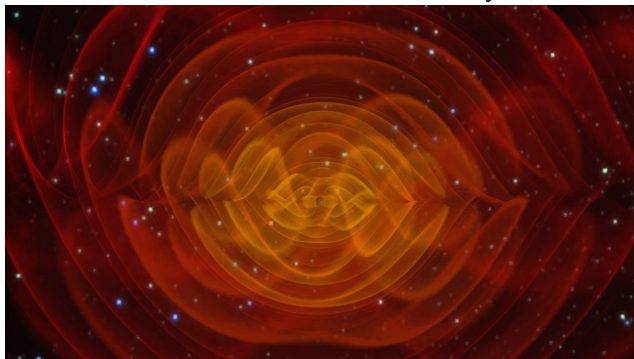
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Amazing achievement of humanity  
from late 1600 till early 2000!

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- ▶ Yes: the dark side of the Universe! (dark matter, dark energy)  
Understanding the dark side may take a couple of decades — interesting times for fundamental physics!

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We may be lucky and witness not only the completion of the Standard Models, but also a first glimpse into the dark side of the Universe within our lifetimes!

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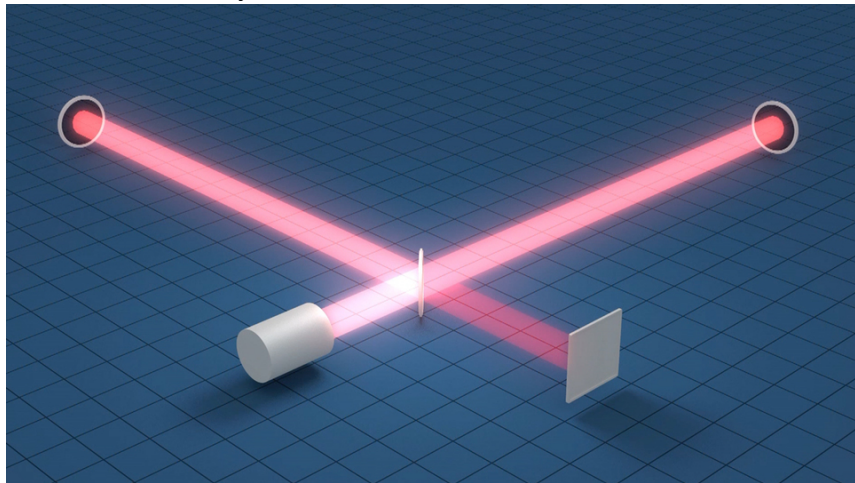
- ▶ In principle easy!
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Typical gravitational waves change spatial distances by a small fraction of the size of a proton!



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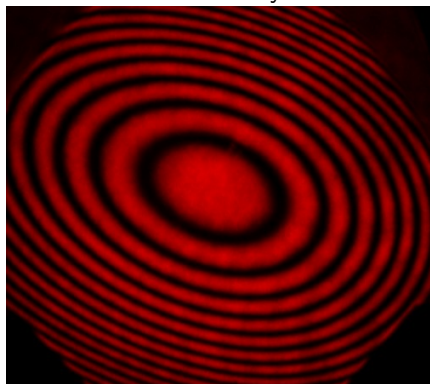
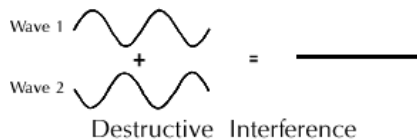
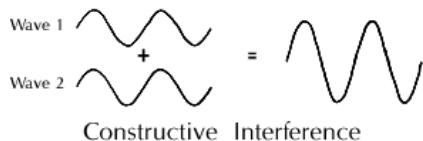
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(Advanced) LIGO needed 25 years of development and 500 million \$ investment, involving more than 900 scientists and engineers

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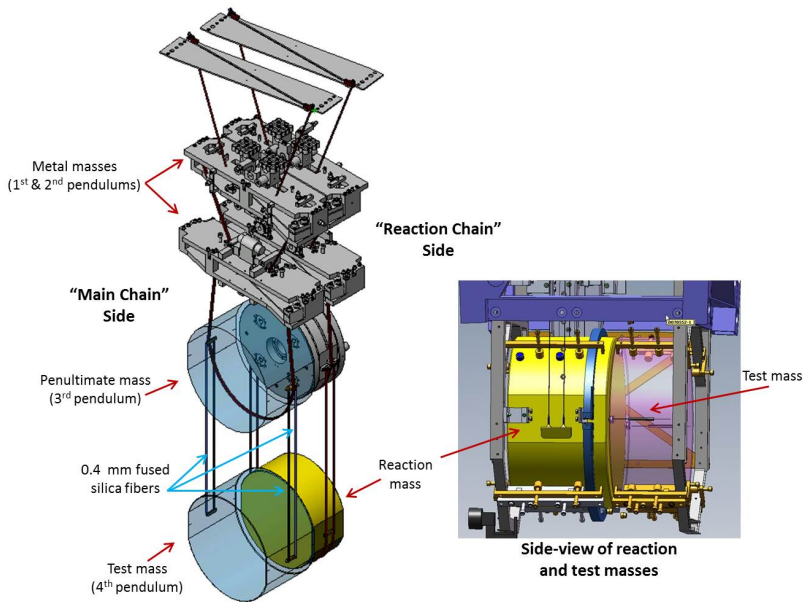


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- ▶ Reduce local errors by suspension system to isolate mirrors from shaking

Suspension system works (in principle) like that of a (very advanced) car — wheels feel bumps, but are decoupled from car

# LIGO suspension system



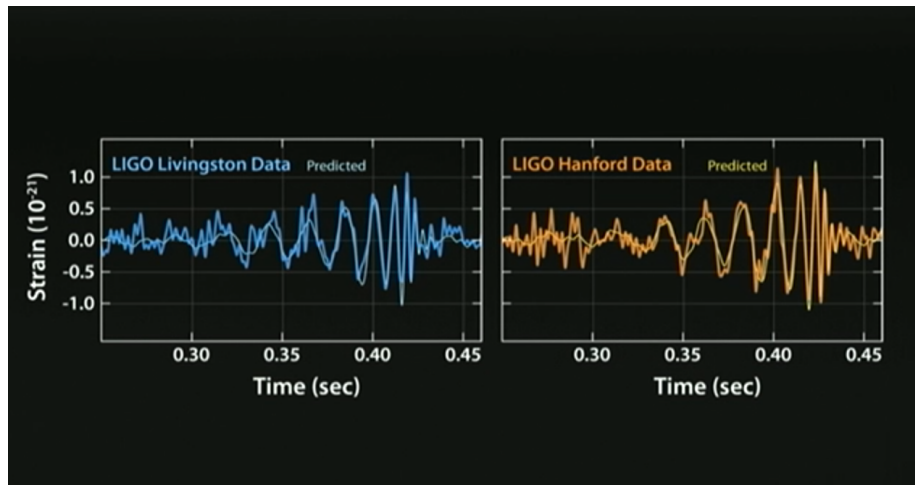
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- ▶ East and West coast data compatible with each other
- ▶ Gravitational wave signal significantly above background
- ▶ Matches very precisely predictions from black hole merger

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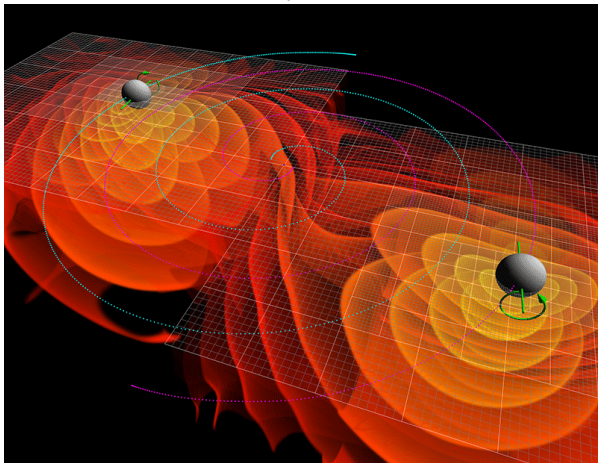
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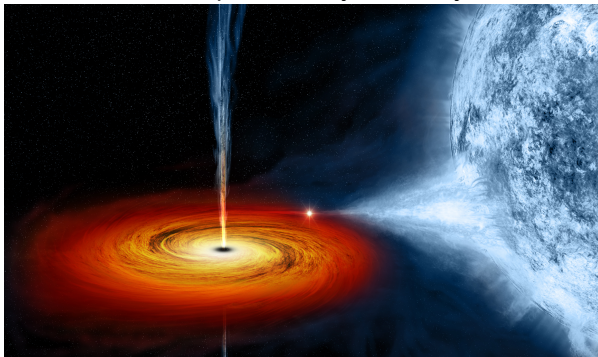


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Having said all this:

- ▶ gravitational waves were expected to exist
- ▶ interest therefore mostly in experimental applications!

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- ▶ Beginning of the era of gravitational wave astronomy
- ▶ You can hide from light in the shadows, but not from gravity

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## Experimental applications: gravitational wave astronomy

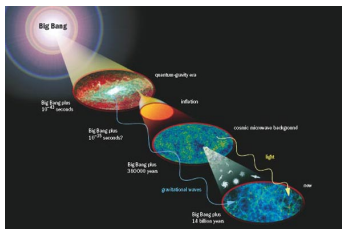
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- ▶ Conversely: Public funding should go to fundamental research, not to industry (Austrian funding agency FWF in dire straits)

Congratulations to the Advanced LIGO team at MIT and 90 other institutions!



I hope you enjoyed my talk!



... any questions?

Backup slide

Educational video by LIGO

