

Black Holes I (136.028)

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Outlook on black hole research, January 2020



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Appetizer, Part I

Physics of the 20th century: harmonic oscillator

Simple idea:

Harmonic oscillator: take a physical system and shake it

Amazingly successful:

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- ▶ QFT corrections to Hydrogen atom



Feynman diagrams contributing to Lamb shift

Appetizer, Part I

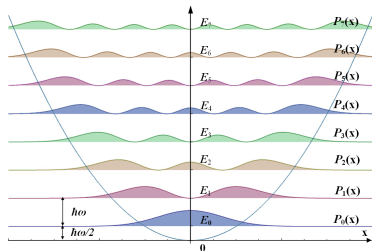
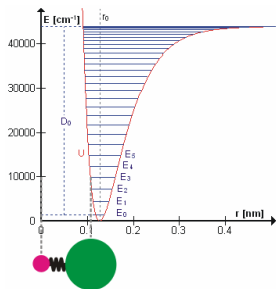
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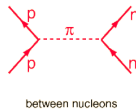
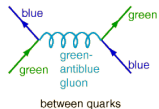
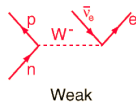
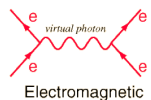
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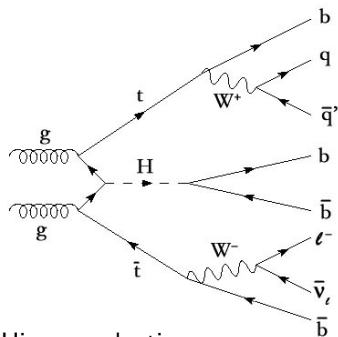
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- ▶ Standard Model of particle physics



Strong Interaction



Higgs production

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Amazingly successful:

- ▶ QFT corrections to Hydrogen atom
- ▶ weakly coupled phonons and electrons in cond-mat
- ▶ Standard Model of particle physics
- ▶ see also the [TU Wien](#) curriculum

Lectures at [TU Wien](#) containing harmonic oscillator (PP = particle physics)

- | | | |
|------------------------|--------------------------|--------------------------|
| ▶ Intro to PP | ▶ Astro-PP | ▶ Path integrals |
| ▶ Atomic, nuclear & PP | ▶ Black holes | ▶ Thermal field theory |
| ▶ Intro to QFT | ▶ Intro to QED | ▶ Cosmology |
| ▶ Theor. methods in PP | ▶ Field theory and pheno | ▶ various basic lectures |

Appetizer, Part II

Physics of the 21st century: black holes?

Application of harmonic oscillator limited to perturbative phenomena

Appetizer, Part II

Physics of the 21st century: black holes?

Application of harmonic oscillator limited to perturbative phenomena

Many physical systems require non-perturbative physics:

- ▶ QCD at low energies
- ▶ High T_c superconductors
- ▶ Graphene
- ▶ Cold atoms
- ▶ Gravity at high curvature

Generally speaking:

Strongly coupled systems require new techniques

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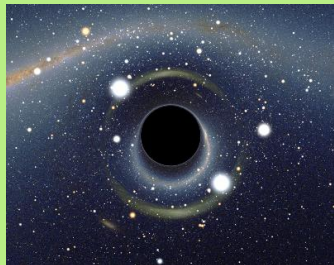
Punch-line of this outlook:

Black hole holography can provide such a technique

Appetizer, Part III

Black holes have apparently paradoxical properties

Black holes: The simplest macroscopic objects in the Universe



Properties determined by:

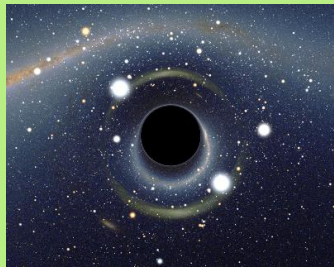
- ▶ Mass M
- ▶ Angular momentum J
- ▶ Charge(s) Q

Black hole \sim elementary particle!

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Black holes: The simplest macroscopic objects in the Universe

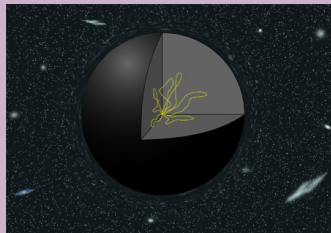


Properties determined by:

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Black hole \sim elementary particle!

Black holes: The most complicated objects conceivable



Quantum mechanics:

- ▶ Black holes radiate
- ▶ Black holes have entropy
- ▶ Black holes are holographic

Bekenstein–Hawking entropy:

$$S_{\text{BH}} \sim A/4$$

Outline

Black hole experiments

Black hole theory

Black hole holography

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Do black holes exist in our Universe? (image source: random webpage)

Ham's scho ans g'sehn? (Ernst Mach to Ludwig Boltzmann concerning the existence of atoms)

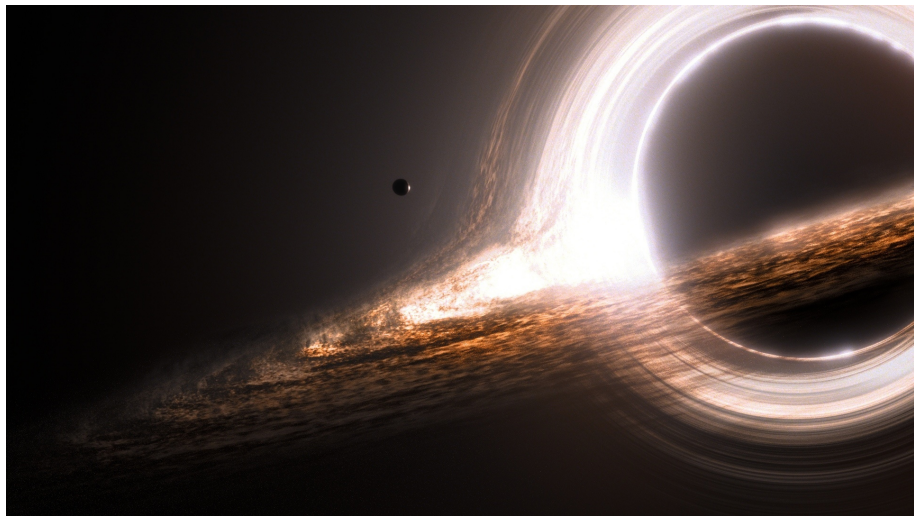
Artistic black hole binary impression



Do black holes exist in our Universe? (image source: [1502.03808](#))

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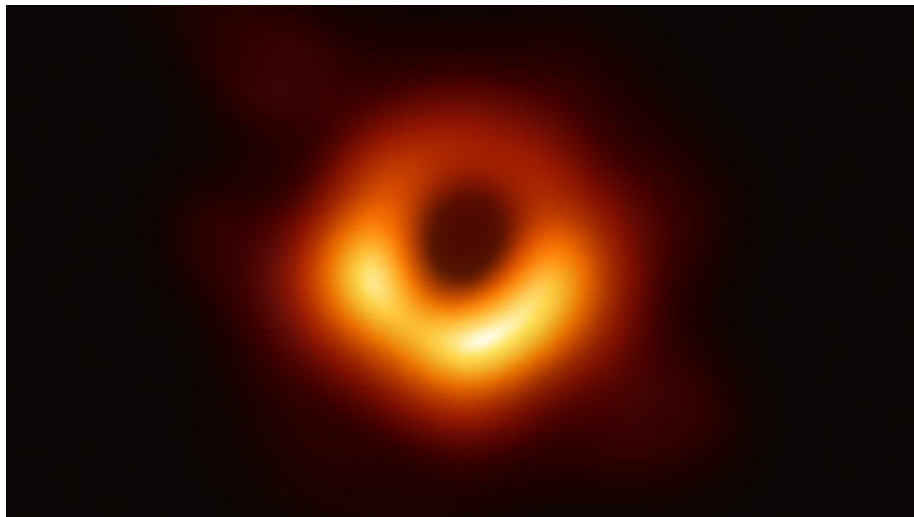
Numerical black hole simulation (interstellar)



Do black holes exist in our Universe? (image source: [1906.11238](#))

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Photo of black hole M87 (EHT collaboration)



Pre-photographic evidence for black holes

“Seeing is believing” is infantile; why did people believe in black holes before the photo?

Experimental evidence for/against various black hole candidates:

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- ▶ stellar black holes: gravitational collapse Chandrasekhar 1930

after fusion in star stops:

Fermi pressure of electrons/neutrons prevents gravitational collapse

critical mass from Fermi pressure = gravitational pressure

rough estimate yields

$$M_{\text{critical}} \sim \frac{1}{m_N^2} \sim 10^{38} \sim 10^{30} \text{ kg} \sim M_{\odot}$$

more refined calculation yields

$$M_{\text{critical}} \approx 3M_{\odot}$$

stellar objects with mass $> 3M_{\odot}$ thus are black holes

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Objects whose mass is clearly beyond critical $M > 3M_{\odot}$ (from '06):

System	P_{orb} [days]	$f(M)$ [M_{\odot}]	Donor Spect. Type	Classification	M_x † [M_{\odot}]
GRS 1915+105 ^a	33.5	9.5 ± 3.0	K/M III	LMXB/Transient	14 ± 4
V404 Cyg	6.471	6.09 ± 0.04	K0 IV	„	12 ± 2
Cyg X-1	5.600	0.244 ± 0.005	09.7 Iab	HMXB/Persistent	10 ± 3
LMC X-1	4.229	0.14 ± 0.05	07 III	„	> 4
XTE J1819-254	2.816	3.13 ± 0.13	B9 III	IMXB/Transient	7.1 ± 0.3
GRO J1655-40	2.620	2.73 ± 0.09	F3/5 IV	„	6.3 ± 0.3
BW Cir ^b	2.545	5.74 ± 0.29	G5 IV	LMXB/Transient	> 7.8
GX 339-4	1.754	5.8 ± 0.5	–	„	„
LMC X-3	1.704	2.3 ± 0.3	B3 V	HMXB/Persistent	7.6 ± 1.3
XTE J1550-564	1.542	6.86 ± 0.71	G8/K8 IV	LMXB/Transient	9.6 ± 1.2
4U 1543-475	1.125	0.25 ± 0.01	A2 V	IMXB/Transient	9.4 ± 1.0
H1705-250	0.520	4.86 ± 0.13	K3/7 V	LMXB/Transient	6 ± 2
GS 1124-684	0.433	3.01 ± 0.15	K3/5 V	„	7.0 ± 0.6
XTE J1859+226 ^c	0.382	7.4 ± 1.1	–	„	„
GS2000+250	0.345	5.01 ± 0.12	K3/7 V	„	7.5 ± 0.3
A0620-003	0.325	2.72 ± 0.06	K4 V	„	11 ± 2
XTE J1650-500	0.321	2.73 ± 0.56	K4 V	„	„
GRS 1009-45	0.283	3.17 ± 0.12	K7/M0 V	„	5.2 ± 0.6
GRO J0422+32	0.212	1.19 ± 0.02	M2 V	„	4 ± 1
XTE J1118+480	0.171	6.3 ± 0.2	K5/M0 V	„	6.8 ± 0.4

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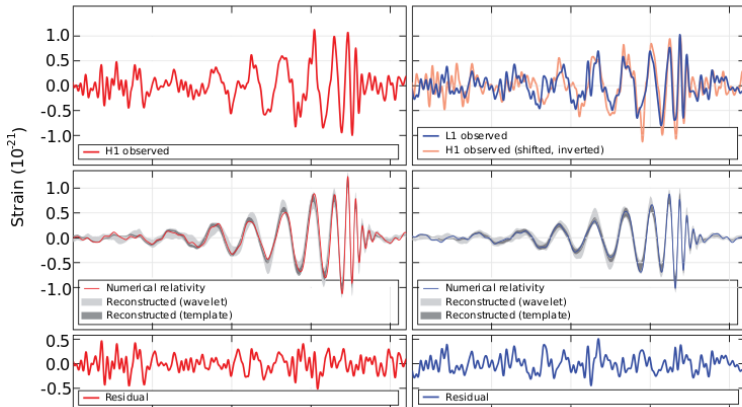
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- ▶ stellar black hole mergers: gravitational wave production **LIGO '16**

Hanford, Washington (H1)

Livingston, Louisiana (L1)

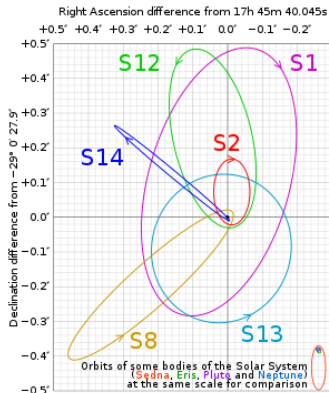
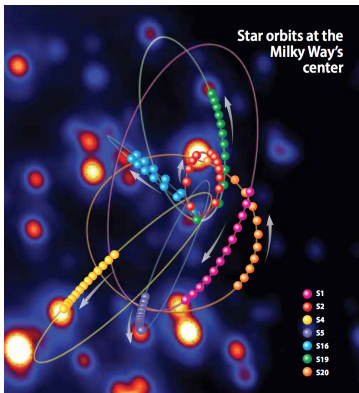


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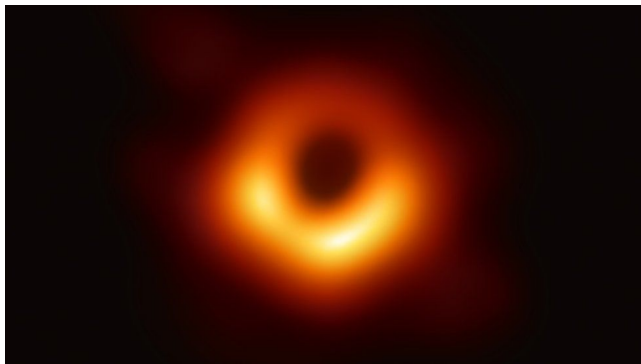


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- ▶ overwhelming evidence for stellar and supermassive black holes
- ▶ confirmed mass ranges so far: $3 - 100 M_{\odot}$ and $10^6 - 10^{10} M_{\odot}$
- ▶ black holes could in principle exist for any mass $> M_{\text{Planck}}$

Outline

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Black hole theory

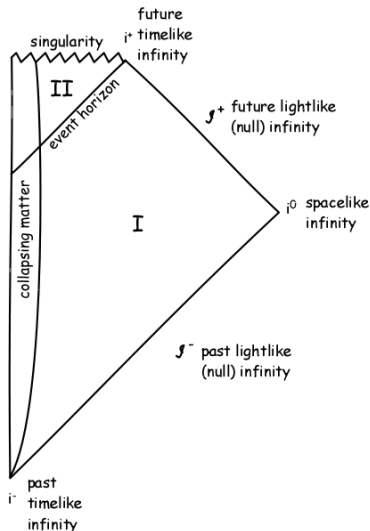
Black hole holography

Definition and observables of black holes

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- ▶ mathematical definition: horizon = boundary of past of null infinity



I outside region

II black hole

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 - ▶ electric monopole mass = mass M
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 - ▶ electric angular momentum = angular momentum J
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Essentially same observables as for elementary particles:
charge Q , mass M , spin $J \Rightarrow$ amazingly simple!

Thermodynamics and black holes — black hole thermodynamics?

Thermodynamics

Zeroth law:

$T = \text{const.}$ in equilibrium

T : temperature

Black hole mechanics

Zeroth law:

$\kappa = \text{const.}$ f. stationary black holes

κ : surface gravity

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Third law:

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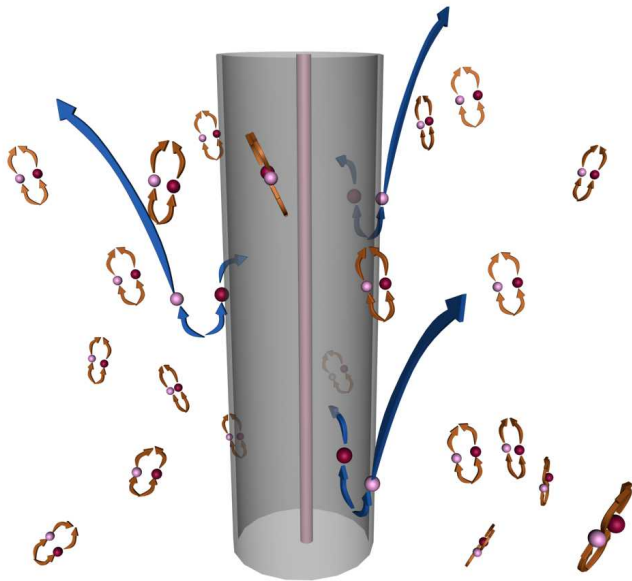
M : mass

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Formal analogy or actual physics?

Hawking effect (QFT under external conditions)

Black holes evaporate due to semi-classical effects!



General black holes:

$$T_{\text{H}} = \frac{\kappa}{2\pi}$$

$$S_{\text{BH}} = \frac{A}{4}$$

Schwarzschild (SI units):

$$T_{\text{H}} = \frac{\hbar c^3}{8\pi G_N k_B M}$$

$$S_{\text{BH}} = \frac{c^3 A}{4G_N \hbar}$$

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Bekenstein–Hawking entropy

Currently template for experimental data in quantum gravity

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$$S_{\text{BH}} = \frac{A}{4} \sim \text{length}^{d-1}$$

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Holographic principle 't Hooft '93; Susskind '95

Gravity in d spatial dimensions equivalent to QFT in $d - 1$ spatial dimensions

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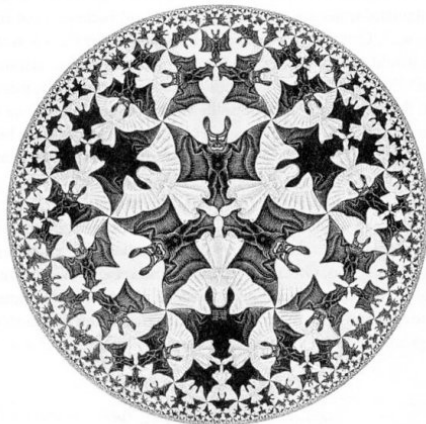
WTF? How can this be true?

Anti-de Sitter/conformal field theory (AdS/CFT) heuristics

Maldacena's [hep-th/9711200](https://arxiv.org/abs/hep-th/9711200) currently has about 15-thousand citations

Best studied realization of holography is AdS/CFT correspondence:

- ▶ AdS is a negatively curved spacetime (maximally symmetric)



Open Universe Looking from inside, boundary at infinity
Limit Circle IV, by M. C. Escher

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$$ds^2 = (E/L)^2 \eta_{\mu\nu} dx^\mu dx^\nu + (L/E)^2 dE^2$$

L sets physical scales

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L sets physical scales and is called “AdS-radius”

This is precisely the line element of AdS in 1 dimension higher!

AdS₃/CFT₂

Focus on symmetries

2d CFTs allow to apply powerful methods; let us focus on CFT₂

- ▶ symmetries of CFT_{*d*}: $SO(d, 2)$ if $d > 2$
- ▶ symmetries of AdS_{*d*+1}: $SO(d, 2)$ if $d > 1$

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- ▶ matches for $d > 2$, but what about $d = 2$?
- ▶ CFT₂ has infinitely many symmetries!

$$[L_n, L_m] = (n - m) L_{n+m} + \frac{c}{12} (n^3 - n) \delta_{n+m, 0} \quad n, m \in \mathbb{Z}$$

symmetry algebra is called 'Virasoro algebra with central charge c '

fineprint: there are actually two copies of this algebra; to reduce clutter we focus on one of them

AdS₃/CFT₂

Focus on symmetries

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at most $D(D + 1)/2$ Killing vectors ξ in D spacetime dimensions

$$\mathcal{L}_\xi g_{\mu\nu} = \xi^\alpha \partial_\alpha g_{\mu\nu} + g_{\mu\alpha} \partial_\nu \xi^\alpha + g_{\nu\alpha} \partial_\mu \xi^\alpha = 0$$

e.g. in $D = 4$ for $g_{\mu\nu} = \eta_{\mu\nu}$ get $\xi^\alpha = \xi_{(0)}^\alpha + \Lambda^\alpha{}_\beta x^\beta = \text{Poincaré trafo}$

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Last question was resolved already in 1960ies by
Bondi, van der Burgh, Metzner and Sachs (BMS)

Asymptotic symmetries

The limit of general relativity at small curvature is special relativity. Right?

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Two important lessons

1. **BMS** symmetries of relevance for QFTs (soft theorems)
2. more generally, gravity theories can have infinitely many symmetries

Back to AdS₃/CFT₂

- ▶ consider asymptotically AdS₃ metrics

$$g_{\mu\nu} = \bar{g}_{\mu\nu}^{\text{AdS}} + \delta g_{\mu\nu}$$

with suitable* choice of $\delta g_{\mu\nu}$

* if you care about details read [Max Riegler's DKPI PhD thesis 1609.02733](#)

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recall CFT₂ symmetries

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Almost Virasoro, so almost there — **but central charge c missing**

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- ▶ **matches perfectly with symmetries of CFT₂!**

Consequences of symmetry matching

What we have learned so far about AdS_3 is already impressive:

3d Einstein gravity with negative cosmological constant and **Brown–Henneaux** boundary conditions, if it exists, is dual to a CFT_2 since the physical Hilbert space falls into representations of two Virasoro algebras

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- ▶ CFT_2 correlation functions calculable on gravity side
e.g. 5-point function of stress-tensor flux components

$$\text{CFT}_2 : \quad \langle T_{++}(z_1)T_{++}(z_2)T_{++}(z_3)T_{++}(z_4)T_{++}(z_5) \rangle = \frac{4c g_5(\gamma, \zeta)}{\prod_{1 \leq i < j \leq 5} z_{ij}}$$

$\gamma = z_{12}z_{34}/(z_{13}z_{24})$, $\zeta = z_{25}z_{34}/(z_{35}z_{24})$, $z_{ij} = z_i - z_j$ and

$$g_5(\gamma, \zeta) = \frac{\gamma + \zeta}{2(\gamma - \zeta)} - \frac{\gamma^2 - \gamma\zeta + \zeta^2}{\gamma(\gamma - 1)\zeta(\zeta - 1)(\gamma - \zeta)} \left([\gamma(\gamma - 1) + 1][\zeta(\zeta - 1) + 1] - \gamma\zeta \right)$$

matches precisely with gravity side

$$\frac{\delta^5 I_{\text{EH}}[g_{\mu\nu}]}{\delta g^{++}(z_1)\delta g^{++}(z_2)\delta g^{++}(z_3)\delta g^{++}(z_4)\delta g^{++}(z_5)} = \frac{4c g_5(\gamma, \zeta)}{\prod_{1 \leq i < j \leq 5} z_{ij}}$$

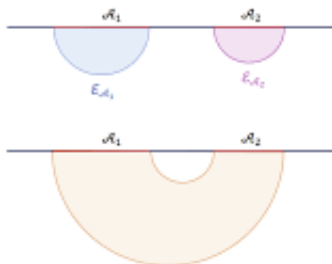
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Holographic dictionary between AdS_3 and CFT_2 observables!

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example: information loss problem in quantum gravity mapped to information gain problem on CFT side

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example: calculation of shear viscosity over entropy density at strong coupling mapped to gravitational wave absorption of black hole

$$\frac{\eta}{s} = \frac{1}{4\pi} + \text{finite coupling corrections}$$

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Black hole holography currently seems like a big hammer!

Black hole holography harmonic oscillator of 21st century?

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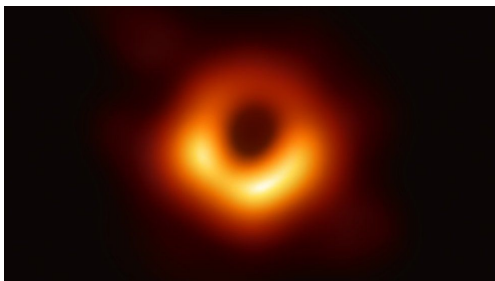
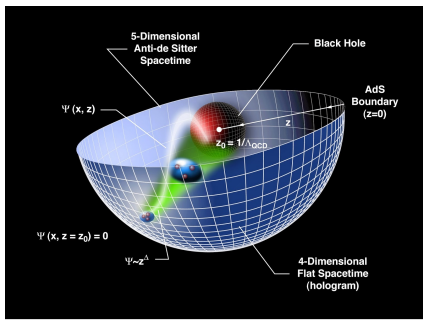
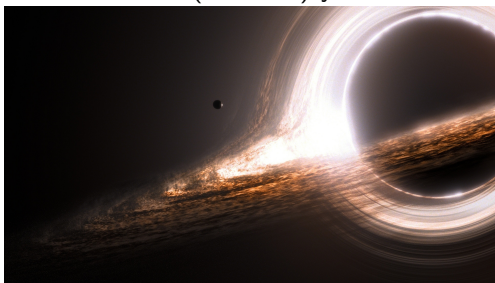
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- ▶ Will need more research on black holes to resolve all issues!
- ▶ Black holes II prepares you for some of the research directions

Thanks for your attention...



...see (some of) you in March!



EHT collaboration, April 2019