

Can Chern-Simons or Rarita-Schwinger be a Volkov-Akulov Goldstone?

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Motivation

- How to break higher-spin symmetries?
- Are they broken spontaneously? What is the mechanism?
- Can one construct consistent non-linear models of higher-spin Goldstone fields associated with spontaneous breaking of HS symmetries?
- To start with a simplified set-up based on simple but yet non-trivial finite HS algebras (similar to SUSY) and see....

History

- In 1975 Hietarinta constructed (super)algebras that generalize to higher-spins conventional susy algebras and proposed to use their non-linear realizations for the description of HS Goldstone fields ala Volkov-Akulov
- Independently, in 1977 Baaklini considered the case of spin-3/2 superalgebra of this type in D=4. 89'-10' Shima, Tanii, Tsuda, Pilot, Rajpoot...
- In 1984 Aragon & Deser constructed, in D=3, "Hypergravity" a system of (non-dynamical) gravity and fermionic spin-(2n+3)/2 gauge field invariant under local spin-(2n+1)/2 supersymmetry transformations
- Recently Hypergravity was revisited by Zinoviev '14, Bunster et.al '14, Fuentealba et.al '15, Henneaux et.al '15, ...

Our motivation: study effects of spontaneous symmetry breaking and properties of non-linear Lagrangians for higher-spin goldstone fields

Simplest cases to be considered: spin-1 and spin-3/2 Goldstones in D=2+1

Hietarinta algebras

Poincaré SUSY algebra

$$\{Q_\alpha, Q_\beta\} = 2\Gamma_{\alpha\beta}^c P_c$$

$$[Q_\alpha, P_c] = 0$$

$$(a = 0, 1, \dots, D-1)$$

α - spinor indices



$$\{Q_\alpha^{a_1 \dots a_n}, Q_\beta^{b_1 \dots b_m}\} = f_{\alpha\beta}^{a_1 \dots a_n, b_1 \dots b_m, c} P_c$$

$$[Q_\alpha^{a_1 \dots a_n}, P_c] = 0$$

Lorentz-invariant
structure constants



Bosonic Hietarinta algebras

$$[S^{a_1 \dots a_n}, S^{b_1 \dots b_m}] = f^{a_1 \dots a_n, b_1 \dots b_m, c} P_c$$

$$[S^{a_1 \dots a_n}, P_c] = 0$$

Difference from the conventional HS algebras:

$$[T_{s_1}, T_{s_2}] = T_{s_1+s_2-2} + T_{s_1+s_2-4} + \dots + T_{|s_1-s_2|+2}$$

(Anti)commutators of the HS generators in the Hietarinta algebras do not produce other yet higher spin generators, i.e. the Hietarinta algebras are finite dimensional.

- Hietarinta algebras can be obtained by contractions of HS algebras

Volkov-Akulov Goldstino models '72

(Describe effects of spontaneous susy breaking in terms of non-linear effective Lagrangians for Goldstone fields on which susy is realized non-linearly)

Poincaré SUSY algebra

$$\{Q_\alpha, Q_\beta\} = 2\Gamma_{\alpha\beta}^c P_c \quad Q_\alpha \leftrightarrow \psi_\alpha(x) \quad \text{Spin } \frac{1}{2} \text{ Goldstino}$$

$$[Q_\alpha, P_c] = 0$$

$$(a = 0, 1, \dots, D-1)$$

VA Construction :

i) Find a susy-invariant Cartan one-form:

$$\Omega = -\mathbf{i}g^{-1}dg = E^a P_a + E^\alpha Q_\alpha, \quad g = e^{ix^a P_a + i\theta^\alpha Q_\alpha}$$

$$E^a = dx^a + i\bar{\theta}\Gamma^a d\theta, \quad E^\alpha = d\theta^\alpha$$

$$\text{SUSY transformations: } \theta'^\alpha = \theta^\alpha + \epsilon^\alpha, \quad x'^a = x^a - \mathbf{i}\bar{\epsilon}\Gamma^a\theta$$

Volkov-Akulov Goldstino models

ii) Replace the Grassmann coordinate with the field: $\theta^\alpha \rightarrow f^{-1}\psi^\alpha(x)$

 susy breaking scale

Susy invariant one-form in space-time:

$$E^a = dx^a + if^{-2}\psi\Gamma^a d\psi = dx^b (\delta_b^a + if^{-2}\psi\Gamma^a \partial_b \psi) \equiv dx^b E_b^a(\psi)$$

susy variations: $\delta x^a = -if^{-2}\varepsilon\Gamma^a\psi, \quad \delta\psi^\alpha = \varepsilon^\alpha + if^{-2}\varepsilon\Gamma^b\psi\partial_b\psi^\alpha$

Non-linear transformation

iii) Construct susy invariant action:

$$S = \frac{f^2}{D!} \int \varepsilon_{a_1 \dots a_D} E^{a_1} \wedge E^{a_2} \dots \wedge E^{a_D} = -f^2 \int d^D x \det E_b^a$$

In $D=3$ $S_{1/2} = \int d^3 x (-i\psi\Gamma^a \partial_a \psi + f^{-2}(\psi\psi)\partial_a \psi\Gamma_b \partial_c \psi \varepsilon^{abc} + \cancel{f^2})$

$$a = 0, 1, 2 \quad \alpha = 1, 2 \quad \psi\psi \equiv \psi^\alpha C_{\alpha\beta}\psi^\beta$$

D=3 Chern-Simons Goldstones

$$[S^a, S^b] = 2i\varepsilon^{abc} P_c$$

$$S^a \rightarrow A_a(x)$$

$$[S^a, P_b] = 0$$

$$\delta A_a(x) = c_a + f^{-2} \varepsilon^{dbc} c_b A_c \partial_d A_a, \quad \delta x^a = -f^{-2} \varepsilon^{dbc} c_b A_c$$

$$[M^a, S^b] = 2i\varepsilon^{abc} S_c$$

↓
constant vector parameter

Invariant one-form: $E^a = dx^a + f^{-2} \varepsilon^{abc} A_b dA_c = dx^d (\delta_d^a + f^{-2} \varepsilon^{abc} A_b \partial_d A_c) \equiv dx^d E_d^a$

Invariant action:

$$S = -f^2 \int d^3x (\det E_d^a - 1) = \int d^3x \left(\overbrace{\varepsilon^{abc} A_a \partial_b A_c}^{\text{Chern-Simons}} - \frac{f^{-2}}{2} \varepsilon^{abc} \varepsilon^{dfg} A_a A_d \partial_b A_f \partial_c A_g \right)$$

Does the action have a gauge symmetry? $\delta A_a = \partial_a \lambda(x) + \dots$

NO \rightarrow vector field acquires a propagating degree of freedom

Propagating CS Goldstone is Galileon

Stueckelberg trick - make the action gauge invariant by replacing

$$A_a \rightarrow \overset{\vee}{A}_a = A_a - f^{1/2} \partial_a \varphi(x), \quad \delta A_a = \partial_a \lambda, \quad \delta \varphi = f^{-1/2} \lambda$$

Make this substitution in the action and consider a so-called decoupling limit $f \rightarrow \infty$

$$\mathcal{L}|_{f \rightarrow \infty} = \varepsilon^{abc} A_a \partial_b A_c - \frac{1}{2} \varepsilon^{abc} \varepsilon^{def} \partial_a \varphi \partial_d \varphi \partial_e \partial_b \varphi \partial_f \partial_c \varphi.$$

Quartic Galileon Lagrangian (appears in modified gravities):

$$\begin{aligned} \mathcal{L}(\varphi) &= \frac{1}{2} \varphi \varepsilon^{abc} \varepsilon^{def} \partial_a \partial_d \varphi \partial_e \partial_b \varphi \partial_f \partial_c \varphi \\ &= -3\varphi \det(\partial_a \partial^b \varphi) \\ &= -\frac{1}{2} \varphi \left((\square \varphi)^3 - 3 \square \varphi \partial_a \partial^b \varphi \partial_b \partial^a \varphi + 2 \partial_a \partial^b \varphi \partial_b \partial^c \varphi \partial_c \partial^a \varphi \right) \end{aligned}$$

Galilean symmetry: $\delta \varphi = c + c_a x^a$

Consequence: Galileon Lagrangians are **quadratic in time derivatives**

Hamiltonian analysis

$$S = \int d^3x \left(\varepsilon^{abc} A_a \partial_b A_c - \frac{f^{-2}}{2} \varepsilon^{abc} \varepsilon^{dfg} A_a A_d \partial_b A_f \partial_c A_g \right)$$

The action is of the first order in time derivative $\partial_0 A_a$

Conjugate momenta are constrained:

$$p^i = \varepsilon^{ij} A_j + f^{-2} F^i(A_b, \partial_i A_c) \quad i = 1, 2, \quad \varepsilon^{ij} \partial_i A_j + O(f^{-2}) \approx 0 \quad \text{secondary constraint}$$
$$p^0 = f^{-2} F^i(A_b, \partial_i A_c)$$

The 4 constraints are of the second class and reduce the number of dofs to 1

On the constraint surface Hamiltonian is

$$H = 6 f^{-2} A_0^2 \det \partial_i A_j, \quad H_\varphi^{\text{Galileon}} = 6 (\partial_0 \varphi)^2 \det \partial_i \partial_j \varphi$$

Hamiltonian is not bounded from below.

Fluctuations around certain zero-energy configurations may have negative energies leading to instabilities

3d Rarita-Schwinger goldstino model

Spin-3/2 Superalgebra

$$\{Q_\alpha^a, Q_\beta^b\} = 2 C_{\alpha\beta} \varepsilon^{abc} P_c + \mathbf{b} \Gamma_{\alpha\beta}^{(a} P^{b)} + \mathbf{c} \eta^{ab} \Gamma_{\alpha\beta}^c P_c, \quad [Q_\alpha^a, P_b] = 0$$

Special cases:

$$\mathbf{b}=4, \mathbf{c}=-2 \quad Q_\alpha \equiv (\Gamma_a Q^a)_\alpha, \quad \hat{Q}^a = Q^a - \frac{1}{3} \Gamma^a Q, \quad \Gamma_a \hat{Q}^a \equiv 0$$

$$\{Q_\alpha, Q_\beta\} = 2 \Gamma_{\alpha\beta}^a P_a, \quad \{Q_\alpha, \hat{Q}_\beta^a\} = 0 = \{\hat{Q}_\alpha^a, \hat{Q}_\beta^b\}$$

$$\mathbf{b}=-4/5, \mathbf{c}=8/5 \quad \{Q_\alpha, \hat{Q}_\beta^a\} = 0 = \{Q_\alpha, Q_\beta\}, \quad \{\hat{Q}_\alpha^a, \hat{Q}_\beta^b\} \neq 0$$

$$\underline{\mathbf{b}=\mathbf{c}=0} \quad \{Q_\alpha^a, Q_\beta^b\} = 2 C_{\alpha\beta} \varepsilon^{abc} P_c, \quad [Q_\alpha^a, P_b] = 0$$

The case of our interest

3d Rarita-Schwinger goldstino model

Spin-3/2 Superalgebra $\{Q_\alpha^a, Q_\beta^b\} = 2C_{\alpha\beta}\epsilon^{abc}P_c, \quad [Q_\alpha^a, P_b] = 0 \quad \alpha, \beta=1,2$

$$Q_\alpha^a \leftrightarrow \chi_\alpha^a(x), \quad \delta\chi_\alpha^a(x) = \zeta_\alpha^a + i f^{-2} \epsilon^{dbc} (\zeta_b^\beta \chi_{c\beta}) \partial_d \chi_\alpha^a,$$

Susy invariant one-form $E^a = dx^a + i f^{-2} \epsilon^{abc} \chi_b^\alpha d\chi_{c\alpha}$

Action $S = \int d^3x \left(\overbrace{i\epsilon^{abc} \chi_a \partial_b \chi_c}^{\text{Rarita-Schwinger}} + \frac{f^{-2}}{2} (\epsilon^{abc} \epsilon^{dfg} - \epsilon^{afc} \epsilon^{dbg}) \chi_a \partial_b \chi_c \chi_d \partial_f \chi_g \right. \\ \left. + \frac{if^{-4}}{6} (\epsilon^{abc} \epsilon^{dfg} - \epsilon^{abg} \epsilon^{dfc}) \epsilon^{pqr} \chi_c \partial_p \chi_g \chi_a \partial_q \chi_b \chi_d \partial_r \chi_f \right)$

Gauge symmetry:

$$\delta\chi_\alpha^a + \frac{if^{-2}}{3} \epsilon^{dfg} \partial_a (\chi_d^\alpha (\chi_f \delta\chi_g)) + i f^{-2} \epsilon^{dfg} (\delta\chi_d \partial_a \chi_f) \chi_g^\alpha = \partial_a \epsilon^\alpha(x)$$

Reduction of the non-linear 3/2-goldstino action to the quadratic RS action

$$S_{\text{RS}} = i \int d^3x \varepsilon^{abc} \hat{\chi}_a \partial_b \hat{\chi}_c, \quad \delta \hat{\chi}_a^\alpha = \partial_a \epsilon^\alpha(x), \quad \text{linearized local «susy»}$$

where
$$\hat{\chi}_a^\alpha = \chi_a^\alpha + \frac{if^{-2}}{3} \varepsilon^{dfg} \chi_d^\alpha (\chi_f \partial_a \chi_g)$$

Non-linear realization of the rigid spin-3/2 susy:

$$\delta \chi_a^\alpha(x) = \zeta_a^\alpha + i f^{-2} \varepsilon^{dbc} (\zeta_b^\beta \chi_{c\beta}) \partial_d \chi_a^\alpha,$$

$$\delta \hat{\chi}_a^\alpha = \zeta_a^\alpha + i f^{-2} \varepsilon^{dbc} (\zeta_b \hat{\chi}_c) \partial_d \hat{\chi}_a^\alpha + \frac{if^{-2}}{3} \varepsilon^{dbc} \left((\hat{\chi}_b \partial_a \hat{\chi}_c) \zeta_d^\alpha + (\zeta_b \partial_a \hat{\chi}_c) \hat{\chi}_d^\alpha \right) + \mathcal{O}(f^{-4})$$

$$[\delta_2, \delta_1] \hat{\chi}_a^\alpha = \xi^d \partial_d \hat{\chi}_a^\alpha, \quad \xi^d = 2 i f^{-2} \varepsilon^{dbc} \zeta_b^1 \zeta_c^2 \quad - \text{ closes on translation}$$

spin-3/2 goldstino is a pure gauge: $\hat{\chi}_a^\alpha = \partial_a \psi^\alpha$ (on – shell)

Summary and things to do

- The simplest examples of spontaneous breaking of Hietarinta symmetries and corresponding Goldstone models turned out to be peculiar non-linear generalizations of the Chern-Simons and Rarita-Schwinger Lagrangians
- The CS Goldstone propagates a scalar mode which is a Galileon field that appears in modified theories of gravity
- ✓ It would be of interest to consider coupling of the CS Goldstone to a **3d bi-gravity** theory which is invariant under local symmetry associated with the algebra

$$\begin{aligned} [S^a, S^b] &= 2i\epsilon^{abc} P_c, & [S^a, P^b] &= [P^a, P^b] = 0 & \quad & 2 \text{ spin-2 gauge fields} \\ [M^a, S^b] &= 2i\epsilon^{abc} S_c, & [M^a, P^b] &= 2i\epsilon^{abc} P_c & \rightarrow & \text{spin-connection} \\ [M^a, M^b] &= 2i\epsilon^{abc} M_c \end{aligned}$$

“Duality” relation with the 3d Maxwell algebra:

$$S^a \leftrightarrow P^a : \quad [P^a, P^b] = 2i\epsilon^{abc} S_c \equiv iS^{ab}, \quad [S^a, S^b] = 0 = [P^a, S^b]$$

3d “Maxwell-Chern-Simons” gravity

P. Salgado, R.J. Szabo and O. Valdivia '14

L. Avilés, E. Frodden, J. Gomis, D. Hidalgo and J. Zanelli '18

$$P^a \leftrightarrow e^a(x) = dx^m e_m^a(x), \quad M^a \leftrightarrow \omega^a(x) = dx^m \omega_m^a(x),$$

$$S^a \leftrightarrow f^a(x) = dx^m f_m^a(x),$$

Local symmetries: $\delta e^a = D\xi^a(x) + \epsilon^{abc} e_b \lambda_c(x) + \varepsilon^{abc} f_b \zeta_c(x)$

$$\delta f^a = D\zeta^a(x) + \epsilon^{abc} f_b \lambda_c(x), \quad \delta \omega^a = D\lambda^a(x)$$

$$S = \mathbf{a}_1 \int (e^a \wedge R_a + \frac{1}{2} f^a \wedge Df_a) + \mathbf{a}_2 \int f^a \wedge R_a, \\ + \mathbf{a}_3 \int (\omega^a d\omega_a + \frac{1}{3} \epsilon_{abc} \omega^a \omega^b \omega^c)$$

$$R^a = d\omega^a + \frac{1}{2} \epsilon^{abc} \omega_b \wedge \omega_c, \quad Df_a = df_a + \varepsilon_{abc} \omega^b \wedge f^c$$

Higher spin generalization $s=n+1$

$$[S^{a_1 \dots a_n}, S^{b_1 \dots b_n}] = 2i \epsilon^{a_1 b_1 c} P_c \eta^{a_2 b_2} \dots \eta^{a_n b_n}, \quad [S^{(a)}, P^b] = [P^a, P^b] = 0$$

$$\begin{aligned} S_{n+1} &= a_1 \int (e^a \wedge R_a + \frac{1}{2} f^{a_1 \dots a_n} \wedge D f_{a_1 \dots a_n}), \\ &+ a_2 \int (\omega^a d\omega_a + \frac{1}{3} \epsilon_{abc} \omega^a \omega^b \omega^c) \end{aligned}$$

Can this action be obtained by a contraction of the conventional CS HS theory based on $SL(n, \mathbb{R}) \times SL(2, \mathbb{R})$?

Conclusion and outlook

- In the spin-3/2 case the spontaneous breaking of the rigid spin-3/2 susy retains local symmetry of the Rarita-Schwinger Lagrangian. **The non-linear spin-3/2 goldstino Lagrangian reduces to the quadratic RS Lagrangian upon a non-linear field redefinition.** This causes the 3d RS goldstino be non-propagating field
- Problems to study:
 - ✓ To couple the RS goldstino to conventional 3d (super)gravity and Hypergravity (with spin-2 and spin-5/2 gauge fields) and to study properties of these models
 - ✓ To consider the RS goldstino model in D=4 associated with the algebra

$$\{Q_\alpha^a, Q_\beta^b\} = 2\varepsilon^{abcd}(\Gamma_5 \Gamma_c)_{\alpha\beta} P_d \quad \alpha, \beta = 1, \dots, 4$$

- ✓ Generalization to higher-spin Goldstone fields