

BMS symmetry in light-cone gravity: A null-front Hamiltonian study

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with Sudarshan Ananth and Lars Brink

Based on arXiv:[2012.07880](#) and [2101.00019](#)

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BMS symmetry at a glance

Bondi approach:

BMS as the asymptotic symmetry group at null infinity

- infinite-dimensional enhancement of Poincaré group
- further extended to include superrotations, $Diff(\mathbb{S}^2)$, near-horizon symmetries
- connections to soft theorems and on-shell amplitudes, Celestial Holography etc.

[**Bondi-van der Burg-Metzner-Sachs '62**, Barnich-Troessaert, Hawking-Perry-Strominger Compère, Detournay, Grumiller, Sheikh-Jabbari, Campiglia-Laddha, Donnay, Zwickel and many more]

Conformal Carroll approach:

- BMS group as conformal Carroll group
- Further extensions to other Carrollian structures
- Symmetries of null hypersurfaces, Carrollian field theory

[**Duval-Gibbons-Hovarth '14**, Ciambelli, Freidel, Flanagan, Leigh, Obers, Petropoulos, ...]

Hamiltonian approach:

BMS symmetry at spatial infinity

- Based on “3 + 1” Hamiltonian formulation of gravity à la ADM
- Canonical realization of the BMS algebra from an action principle
- a precursor to any quantization methods

[**Henneaux-Troessaert '18**, Fuentealba, Matulich, Tanzi, Guilini, ...]

The other Hamiltonian formulation of gravity

“Forms of relativistic dynamics” [Dirac '49] → Use a null time parameter to study dynamics

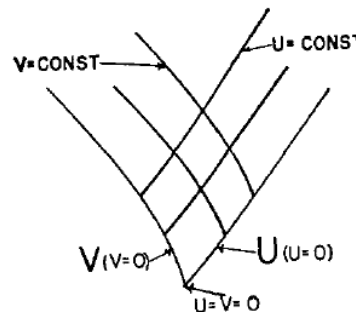
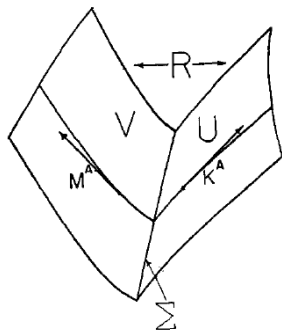
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“Covariant 2+2 formulation of the initial-value problem in general relativity”

[d'Inverno and Smallwood '79] [Gambini-Restuccia, C. Torre, M. Kaku,...]



- Spacelike foliation of codim 2 (instead of 1)
- Unconstrained Hamiltonian systems
- Gravitational d.o.f. identified with the “conformal two-metric”

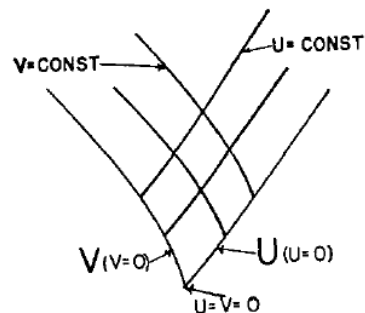
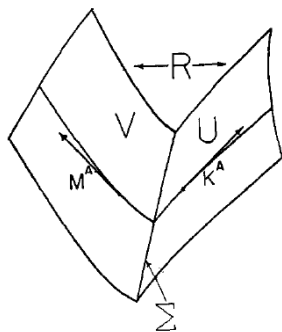
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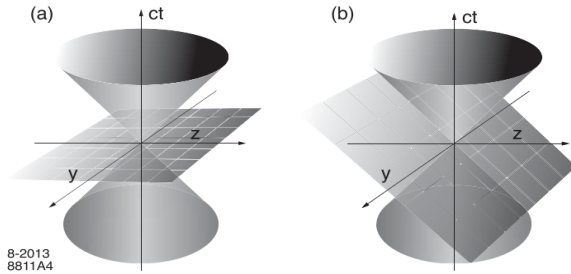


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Focus of this talk: BMS symmetry in the other Hamiltonian formulation

- Set up the 2+2 Hamiltonian formulation of light-cone gravity
- Study the BMS symmetry from residual gauge invariance

Light-cone coordinates: The front-form



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“Forms of relativistic dynamics” [Dirac '49]

(a) Instant form: time x^0

Initial data on a spatial hyperplane ($x^0 = 0$)

(b) Front form: time $x^+ = \frac{x^0 + x^3}{\sqrt{2}}$

Initial data on a null hyperplane ($x^+ = 0$)

- Poincaré generators in the instant form: $(P_\mu, M_{\mu\nu})$

$$[P, P] \sim 0, \quad [P, M] \sim P, \quad [M, M] \sim M$$

$(P_0, M_{0i}) \rightarrow$ four “Hamiltonians”

- Poincaré generators in light-cone coordinates, $x^\mu = (x^+, x^-, x^i)$, $i = 1, 2$

Kinematical $K = \{P_i, P_-, M_{ij}, M_{+-}\}$,

Dynamical $D = \{P_+, M_{i+}\} \rightarrow$ three “Hamiltonians” pick up corrections

Non-linear terms in \mathbb{D} generators give us a handle on the dynamics of the interacting theory

Light-cone Poincaré algebra in $d = 4$

- Non-vanishing commutators of the Poincaré algebra

$$J^+ = \frac{J^{+1} + iJ^{+2}}{\sqrt{2}}, \quad \bar{J}^+ = \frac{J^{+1} - iJ^{+2}}{\sqrt{2}}, \quad J = J^{12}, \quad H = P_+ = -P^-.$$

$$\begin{aligned} [H, J^{+-}] &= -iH, & [H, J^+] &= -iP, & [H, \bar{J}^+] &= -i\bar{P} \\ [P^+, J^{+-}] &= iP^+, & [P^+, J^-] &= -iP, & [P^+, \bar{J}^-] &= -i\bar{P} \\ [P, \bar{J}^-] &= -iH, & [P, \bar{J}^+] &= -iP^+, & [P, J] &= P \end{aligned}$$

... and many more

[Bengtsson-Bengtsson-Brink '83]

- Underlying Carrollian algebra

Rotation $\mathbb{J} = \{J^{12}, J^{+-}, J^+, \bar{J}^+\}$, Boosts $\mathbb{K} = \{J^-, \bar{J}^-\}$

Translations $\mathbb{P} = \{P, \bar{P}, P_-\}$, Hamiltonian $\mathbb{H} = P_+$

$$\begin{aligned} [\mathbb{J}, \mathbb{J}] &= \mathbb{J}, & [\mathbb{J}, \mathbb{P}] &= \mathbb{P}, & [\mathbb{J}, \mathbb{K}] &= \mathbb{K} \\ [\mathbb{J}, \mathbb{H}] &= 0, & [\mathbb{H}, \mathbb{P}] &= 0, & [\mathbb{H}, \mathbb{K}] &= 0 \\ [\mathbb{P}, \mathbb{P}] &= 0, & [\mathbb{K}, \mathbb{K}] &= 0, & [\mathbb{P}, \mathbb{K}] &= \mathbb{H} \end{aligned}$$

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- In terms of the **Kinematical-Dynamical split**

$$\mathbb{K} = \{P_i, P_-, M_{ij}, M_{+-}\}, \quad \mathbb{D} = \{P_+, M_{i+}\}$$

$$[\mathbb{K}, \mathbb{K}] = \mathbb{K}, \quad [\mathbb{K}, \mathbb{D}] = \mathbb{D}, \quad [\mathbb{D}, \mathbb{D}] = 0$$

Gravity in the light-cone gauge

“ l_{C2} formalism” [Scherk-Schwarz, Schwarz-Goroff, Bengtsson-Cederwall-Lindgren]

- Einstein-Hilbert action in 4D

$$\mathcal{S} = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} R[g_{\mu\nu}]$$

- Light-cone gauge: Set the “minus” components to zero

$$g_{--} = g_{-i} = 0, \quad (i = 1, 2)$$

$$10 - 3 = 7$$

Parametrization

$$g_{+-} = -e^\phi, \quad g_{ij} = e^\psi \gamma_{ij}$$

ϕ, ψ, γ_{ij} are real and $\det \gamma_{ij} = 1$

Light-cone gravity metric

$$dS_{LC}^2 = g_{\mu\nu} dx^\mu dx^\nu = -2e^\phi dx^+ dx^+ + g_{++} (dx^+)^2 + g_{+i} dx^+ dx^i + e^\psi \gamma_{ij} dx^i dx^j$$

given in terms of 7 functions $\{\phi, \psi, \gamma_{ij}, g_{++}, g_{+i}\}$

- “2+2” split of the Einstein field equations $R_{\mu\nu} = 0$ [Sachs, d’Inverno-Smallwood, ...]

Dynamical equations: $R_{ij} = 0$

Constraint equations: $R_{--} = R_{-i} = 0$

Subsidiary equations: $R_{++} = R_{+i} = 0$

Trivial equations: $R_{+-} = 0$

Gravity in the light-cone gauge

Can we solve the constraint equations?

→ subject to choice of coordinates, gauge conditions and iterative integration schemes

- Constraint equation $R_{--} = 0$

$$2\partial_- \phi \partial_- \psi - (\partial_- \psi)^2 - 2\partial_-^2 \psi + \frac{1}{2} \partial_- \gamma^{ij} \partial_- \gamma_{ij} = 0.$$

Fourth gauge choice : [Scherk-Schwarz]

$$\phi = \frac{\psi}{2}$$

$$7 - 1 = 6$$

allows us to integrate [†] out ψ

$$\psi = \frac{1}{4} \frac{1}{\partial_-^2} (\partial_- \gamma^{ij} \partial_- \gamma_{ij})$$

$$6 - 1 = 5$$

[†] Our *integration scheme*

$$\partial_- f(x^-) = g(x^-) \Rightarrow f(x^-) = \frac{1}{\partial_-} g(x^-) = - \int \epsilon(x^- - y^-) g(y^-) dy^- + \text{“constant”}$$

- The constraint $R_{-i} = 0$ eliminates g_{+i}

$$5 - 2 = 3$$

↑
FIXED BY
INITIAL DATA

- $R_{-+} = 0$ allows us to eliminates g_{++}

$$3 - 1 = 2$$

Light-cone action for gravity

- Closed form expression

$$\begin{aligned} S[\gamma_{ij}] = & \frac{1}{2\kappa^2} \int d^4x e^\psi \left(2\partial_+\partial_-\phi + \partial_+\partial_-\psi - \frac{1}{2}\partial_+\gamma^{ij}\partial_-\gamma_{ij} \right) \\ & - e^\phi \gamma^{ij} \left(\partial_i\partial_j\phi + \frac{1}{2}\partial_i\phi\partial_j\phi - \partial_i\phi\partial_j\psi - \frac{1}{4}\partial_i\gamma^{kl}\partial_j\gamma_{kl} + \frac{1}{2}\partial_i\gamma^{kl}\partial_k\gamma_{jl} \right) \\ & - \frac{1}{2}e^{\phi-2\psi}\gamma^{ij}\frac{1}{\partial_-}R_i\frac{1}{\partial_-}R_j, \end{aligned}$$

where

$$R_i \equiv e^\psi \left(\frac{1}{2}\partial_-\gamma^{jk}\partial_i\gamma_{jk} - \partial_-\partial_i\phi - \partial_-\partial_i\psi + \partial_i\phi\partial_-\psi \right) + \partial_k(e^\psi\gamma^{jk}\partial_-\gamma_{ij})$$

Light-cone action for gravity

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- Perturbative expansion

$$\gamma_{ij} = (e^{\kappa H})_{ij}, \quad H = \begin{pmatrix} h_{11} & h_{12} \\ h_{12} & h_{22} \end{pmatrix}; \quad h_{22} = -h_{11}$$

Define

$$h = \frac{1}{\sqrt{2}}(h_{11} + i h_{12}), \quad \bar{h} = \frac{1}{\sqrt{2}}(h_{11} - i h_{12})$$

- Light-cone Lagrangian

$$\mathcal{L} = \frac{1}{2}\bar{h} \square h + 2\kappa \bar{h} \partial_-^2 \left(\frac{\bar{\partial}}{\partial_-} h \frac{\bar{\partial}}{\partial_-} h - h \frac{\bar{\partial}^2}{\partial_-^2} h \right) + c.c. + \text{higher order terms}$$

h and \bar{h} represent gravitons of helicity +2 and -2 respectively

Light-cone Hamiltonian for gravity

- Conjugate momenta (recall: x^+ is time)

$$\mathcal{L} = \bar{h} (\partial_- \partial_+ - \partial \bar{\partial}) h + 2\kappa \bar{h} \partial_-^2 \left(\frac{\bar{\partial}}{\partial_-} h \frac{\bar{\partial}}{\partial_-} h - h \frac{\bar{\partial}^2}{\partial_-^2} h \right) + \dots$$

$$\pi = \frac{\delta \mathcal{L}}{\delta(\partial_+ h)} = -\partial_- \bar{h}, \quad \bar{\pi} = \frac{\delta \mathcal{L}}{\delta(\partial_+ \bar{h})} = -\partial_- h$$

$(\pi, \bar{\pi})$ not independent variables \Rightarrow **Half the d.o.f than in the 3+1 formalism**

\rightarrow a feature of *all* null-front Hamiltonian systems

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- Light-cone Hamiltonian for gravity

$$\mathcal{H} = \partial \bar{h} \bar{\partial} h + 2\kappa \partial_-^2 \bar{h} \left(h \frac{\bar{\partial}^2}{\partial_-^2} h - \frac{\bar{\partial}}{\partial_-} h \frac{\bar{\partial}}{\partial_-} h \right) + c.c. + \mathcal{O}(\kappa^2)$$

- Poisson brackets

$$[h(x), \pi(y)] = \delta(x^- - y^-) \delta^{(2)}(x - y) \Rightarrow [h(x), \bar{h}(y)] = \underbrace{\frac{1}{\partial_-} \delta(x^- - y^-)}_{\text{STEP FUNCTION}} \delta^{(2)}(x - y),$$

$$[h(x), h(y)] = [\bar{h}(x), \bar{h}(y)] = 0.$$

[Scherk-Schwarz' 75, Bengtsson-Cederwall-Lindgren '83]

Notion of symmetry

A canonical transformation $(h, \bar{h}) \xrightarrow{\delta_X} (\tilde{h}, \tilde{\bar{h}})$

which leaves the action invariant

$$\delta_X \mathcal{S}[h, \bar{h}] = 0$$

Transformation laws = P.B. with the generator $G_X[h, \bar{h}]$, e.g.

$$\delta_X h = \{G_X, h\}_{PB}$$

For example,

Poincaré generators in terms of the fields h and \bar{h}

$$H = P_+ = \int d^3x \mathcal{H}(h, \bar{h}), \quad P = \int d^3x \partial_- \bar{h} \partial h, \quad P_- = \int d^3x \partial_- \bar{h} \partial_- h, \quad \dots$$

$$J = i \int d^3x \partial_- \bar{h} (x \bar{\partial} - \bar{x} \partial - 2) h,$$

$$J^- = \int d^3x [x \mathcal{H}(h, \bar{h}) + \partial_- \bar{h} (x^- \partial - 2 \frac{\partial}{\partial_-}) h + \mathcal{S}], \quad \dots$$

→ canonical realization of Poincaré algebra in light-cone gravity

BMS symmetry from residual gauge invariance

Is there any residual reparameterization freedom left?

$$x^\mu \rightarrow x^\mu + \xi^\mu$$

First gauge condition holds

$$g_{--} = 0 \Rightarrow \partial_- \xi^+ = 0 \Rightarrow \xi^+ = f(x^+, x^j)$$

Second gauge condition $g_{-i} = 0$ leads to a relation between ξ^+ and ξ^i

$$\partial_- \xi^j g_{ij} + \partial_i \xi^+ g_{+-} = 0$$

Fourth gauge condition fixes x^+ dependence of $f(x^+, x^j)$

$$\xi^+ = f = \frac{1}{2} x^+ \partial_i Y^i + T(x^k)$$

$$\xi^i = -\partial_k f \frac{1}{\partial_-} (g_{-+} g^{ik}) + Y^i(x^k)$$

$$\xi^- = -\partial_i Y^i x^- + (\partial_+ \xi_i) x^i$$

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ONLY INDEPENDENT
PARAMETER

→ leaves the light-cone gravity action invariant

$$\delta_\xi \mathcal{S}[h, \bar{h}] = 0$$

iff $\partial^2 Y = 0 \rightarrow$ only Lorentz rotations (no supertranslations)

[Ananth, Brink and SM]

BMS algebra in light-cone gravity

- BMS transformation law (on the initial surface $x^+ = 0$),

$$\begin{aligned} \delta_{Y, \bar{Y}, T} h &= Y(x) \bar{\partial} h + \bar{Y}(\bar{x}) \partial h + (\partial \bar{Y} - \bar{\partial} Y) h + T \frac{\partial \bar{\partial}}{\partial_-} h \\ &\quad - 2 \kappa T \partial_- \left(h \frac{\bar{\partial}^2}{\partial_-^2} h - \frac{\bar{\partial}}{\partial_-} h \frac{\bar{\partial}}{\partial_-} h \right) - 2 \kappa T \frac{1}{\partial_-} \left(\frac{\partial^2}{\partial_-^2} \bar{h} \partial_-^2 h \right) \\ &\quad - 2 \kappa T \frac{\partial^2}{\partial_-^3} (\bar{h} \partial_-^2 h) + 4 \kappa T \frac{\partial}{\partial_-^2} \left(\frac{\partial}{\partial_-} \bar{h} \partial_-^2 h \right) + \mathcal{O}(\kappa^2) \end{aligned}$$

- Symmetry algebra

$$\left[\delta(Y_1, \bar{Y}_1, T_1), \delta(Y_2, \bar{Y}_2, T_2) \right] h = \delta(Y_{12}, \bar{Y}_{12}, T_{12}) h,$$

with parameters

$$\begin{aligned} Y_{12} &\equiv Y_2 \bar{\partial} Y_1 - Y_1 \bar{\partial} Y_2 \\ \bar{Y}_{12} &\equiv \bar{Y}_2 \partial \bar{Y}_1 - \bar{Y}_1 \partial \bar{Y}_2 \\ T_{12} &\equiv [Y_2 \bar{\partial} T_1 + \bar{Y}_2 \partial T_1 + \frac{1}{2} T_2 (\bar{\partial} Y_1 + \partial \bar{Y}_1)] - (1 \leftrightarrow 2). \end{aligned}$$

→ BMS algebra in four dimensions

- Canonical generator for supertranslations

$$\begin{aligned} \delta_T h &= [G_T, h], & \delta_T \bar{h} &= [G_T, \bar{h}]. \\ G_T &= \int d^3x \partial_- \bar{h} (\delta_T h) = \int d^3x \partial_- \bar{h} T \frac{\partial \bar{\partial}}{\partial_-} h + \mathcal{O}(\kappa), \end{aligned}$$

Light-cone representation of the BMS algebra

- Light-cone Poincaré algebra

$$\mathbb{K} : \{P, \bar{P}, P^+, J, J^+, \bar{J}^+, J^{+-}\}$$

$$\mathbb{D} : \{P^- \equiv H, J^-, \bar{J}^-\}$$

$$[\mathbb{K}, \mathbb{K}] = \mathbb{K}, \quad [\mathbb{K}, \mathbb{D}] = \mathbb{D}, \quad [\mathbb{D}, \mathbb{D}] = 0.$$

- Light-cone BMS algebra

$$\mathbb{K} \rightarrow \mathbb{K},$$

$$\mathbb{D} \rightarrow \mathbb{D}(T),$$

$$[\mathbb{K}, \mathbb{K}] = \mathbb{K}, \quad [\mathbb{K}, \mathbb{D}(T)] = \mathbb{D}(T), \quad [\mathbb{D}(T), \mathbb{D}(T)] = 0.$$

Dynamical part enhanced to infinite-dimensional supertranslations
labeled by a single parameter T

- Poincaré part of the BMS

$$\partial^2 T = \bar{\partial}^2 T = 0$$

$\Rightarrow \mathbb{D}(T)$ reduces to $\mathbb{D} : \{H, J^-, \bar{J}^-\} \rightarrow$ the three “Hamiltonians” of Dirac

$$\begin{array}{ccc} \swarrow & \downarrow & \searrow \\ a\partial_+ & ibx\partial_+ & -i\bar{b}\bar{x}\partial_+ \end{array}$$

[Ananth, Brink and SM]

Light-cone BMS algebra: key features

- BMS from residual gauge invariance in the bulk: No reference to asymptotic limits
- Dynamical part of the algebra enhanced to infinite-dimensional supertranslations
- Lorentz invariance of the Hamiltonian eliminates superrotations

Some open questions

- How is the light-cone initial data is related to null infinity data?
- Connections to on-shell amplitudes and geometry on null hypersurfaces
- Self-dual and anti self-dual gravity; double copy results

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